# Second-order refraction and diffraction of surface water waves 

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(Received 6 May 2005 and in revised form 20 September 2005)


#### Abstract

The mild-slope equation is an effective approximation for treating the combined effects of refraction and diffraction of infinitesimal water waves, for it reduces the spatial dimension of the linear boundary-value problem from three to two. We extend this approximation to nonlinear waves up to the second order in wave steepness, in order to simplify the inherently three-dimensional task. Assuming that the geometrical complexity is restricted to a finite, though large, horizontal domain, the hybrid-element method designed earlier for linearized problems is modified for the two-dimensional elliptic boundary-value problems at the second order. In the special case of a semicircular peninsula (or a vertical cylinder on a cliff) in a sea of constant depth, the solution is analytical. Effects of the angle of incidence are examined for the freesurface height along the cylinder. For a cylinder standing on a shoal of radially varying depth, numerical results are discussed.


## 1. Introduction

The three-dimensional problem of nonlinear diffraction is difficult even for a sea of constant depth bounded by vertical cliff-like coasts. After several earlier attempts, the second-order theory for an axially symmetric body was finally completed by Kim \& Yue (1989) who employed a semi-analytical method based on boundary elements. The special case of a vertical circular cylinder was later solved analytically by Chau \& Eatock Taylor (1992). Their methods of analysis are fully three-dimensional and have not been extended to nonlinear diffraction and refraction and may be costly if the bathymetry is non-uniform over many wavelengths.

On the other hand, for infinitesimal waves, the combined diffraction and refraction by scatterers over a slowly varying bathymetry can be efficiently analysed by the mild-slope equation (MSE) developed from the linearized theory by Berkhoff (1972). A special advantage of MSE is that it reduces the boundary-value problem from three dimensions to two, and hence facilitates numerical computations involving a large horizontal domain. In the original derivation of Berkhoff (see also Smith \& Sprinks (1975) who derived MSE by weighted averaging via Green's formula) terms proportional to the small bed slope $\nabla h$ were kept, but those proportional to $(\nabla h)^{2}, \nabla^{2} h$ were neglected. For applications to steeper bathymetries, various modifications have been proposed (Kirby 1986; Massel 1993; Chamberlain \& Porter 1995; Porter \& Staziker 1995; Miles \& Chamberlain 1998; Agnon 1999; Athanassoulis \& Belibassakis 1999). In particular, Chamberlain \& Porter retain the second-order depth-gradient terms and show that the modified mild-slope equation has a better accuracy for scattering by a corrugated seabed of length scale comparable to the surface waves.

Increased accuracy can be achieved for still steeper bed slopes by including all the evanescent modes, leading to an infinite set of two-dimensional modified mild-slope equations coupling all vertical modes (Massel 1993; Porter \& Staziker 1995). As far as we are aware, extensions of the mild-slope approximation to nonlinear problems are so far limited to basic ideas, but not yet fully implemented (Agnon 1999).

In coastal dynamics there is a need for the prediction of the diffraction of nonlinear waves over a seabed of depth varying slowly over a large area. Efficient methods that can lighten the computational tasks in three dimensions are required. In this paper, we shall extend the theoretical ideas of the mild-slope approximation to nonlinear diffraction/refraction, so as to simplify the three-dimensional computations. It will be shown that, at second order, the mild-slope equations are a set of twodimensional coupled elliptic and inhomogeneous partial differential equations with variable coefficients. Application will first be made to the limiting case of a semicircular cylinder on a wall in a sea of constant depth, for which the equations become uncoupled and the problem is solved analytically. For cases of varying depth, the hybrid-element method of Chen \& Mei (1974) for linearized scattering of infinitesimal waves is modified for the inhomogeneous boundary-value problems. A numerical example involving radially varying depth is discussed.

## 2. Perturbation equations

We assume the fluid to be incompressible and inviscid, and the flow irrotational. Let the velocity field be represented as the gradient of a scalar potential, i.e. $(u, v, w)=\nabla_{3} \Phi$. Incompressibility requires that

$$
\begin{equation*}
\nabla_{3}^{2} \Phi=\nabla^{2} \Phi+\frac{\partial^{2} \Phi}{\partial z^{2}}=0 \tag{2.1}
\end{equation*}
$$

in the fluid, where $\nabla$ and $\nabla_{3}$ stand for two- and three-dimensional gradient operators, respectively. For small-amplitude waves, i.e. $\epsilon \equiv k A \ll 1$, where $A$ is the amplitude of the incident waves, we expand the free-surface conditions in powers of $\epsilon \equiv k A$ about $z=0$, and obtain the approximate dynamic boundary condition,

$$
\begin{equation*}
g \zeta+\frac{\partial \Phi}{\partial t}+\zeta \frac{\partial^{2} \Phi}{\partial t \partial z}+\frac{1}{2}\left(\nabla_{3} \Phi\right)^{2}=O\left(\epsilon^{3}\right), \quad z=0 \tag{2.2}
\end{equation*}
$$

where $\zeta$ denotes the vertical displacement of the free surface. The preceding equation can be combined with the approximate kinematic condition to give,

$$
\begin{equation*}
\frac{\partial \Phi}{\partial z}+\frac{1}{g} \frac{\partial^{2} \Phi}{\partial t^{2}}=-\frac{\zeta}{g} \frac{\partial}{\partial z}\left(g \frac{\partial \Phi}{\partial z}+\frac{\partial^{2} \Phi}{\partial t^{2}}\right)-\frac{1}{g} \frac{\partial}{\partial t}\left(\nabla_{3} \Phi\right)^{2}+O\left(\epsilon^{3}\right), \quad z=0 \tag{2.3}
\end{equation*}
$$

On the seabed, $z=-h(x, y)$, no flux requires that

$$
\begin{equation*}
\frac{\partial \Phi}{\partial z}=-\nabla \Phi \cdot \nabla h, \quad z=-h(x, y) \tag{2.4}
\end{equation*}
$$

All lateral boundaries $\partial B$, including the coast, are assumed to be vertical, so that

$$
\begin{equation*}
\frac{\partial \Phi}{\partial n}=n \cdot \nabla \Phi=0, \quad(x, y) \in \partial B \tag{2.5}
\end{equation*}
$$

We now introduce the perturbation expansions

$$
\begin{equation*}
\Phi=\epsilon \Phi_{1}+\epsilon^{2} \Phi_{2}+\cdots, \quad \zeta=\epsilon \zeta_{1}+\epsilon^{2} \zeta_{2}+\cdots \tag{2.6}
\end{equation*}
$$

into (2.1) to (2.5). At first and second order, both $\Phi_{1}$ and $\Phi_{2}$ satisfy the Laplace equation in the fluid, (2.4) on the sloping seabed, and (2.5) along all lateral boundaries. On the still-water free surface, $z=0$, the boundary condition for the first-order potential is homogeneous

$$
\begin{equation*}
g \frac{\partial \Phi_{1}}{\partial z}+\frac{\partial^{2} \Phi_{1}}{\partial t^{2}}=0, \quad z=0 \tag{2.7}
\end{equation*}
$$

while that for the second-order potential is not

$$
\begin{equation*}
\frac{\partial \Phi_{2}}{\partial z}+\frac{1}{g} \frac{\partial^{2} \Phi_{2}}{\partial t^{2}}=\frac{1}{g^{2}} \frac{\partial \Phi_{1}}{\partial t} \frac{\partial}{\partial z}\left[g\left(\frac{\partial \Phi_{1}}{\partial z}\right)+\frac{\partial^{2} \Phi_{1}}{\partial t^{2}}\right]-\frac{1}{g} \frac{\partial}{\partial t}\left(\nabla_{3} \Phi_{1}\right)^{2}, \quad z=0 \tag{2.8}
\end{equation*}
$$

The free-surface displacement at first order, $\zeta_{1}$, is related to $\Phi_{1}$ by

$$
\begin{equation*}
\zeta_{1}=-\frac{1}{g}\left[\frac{\partial \Phi_{1}}{\partial t}\right]_{z=0} \tag{2.9}
\end{equation*}
$$

while the second-order correction, $\zeta_{2}$, consists of two parts,

$$
\begin{equation*}
\zeta_{2}=\zeta_{2}^{(1)}+\zeta_{2}^{(2)} \tag{2.10}
\end{equation*}
$$

One part can be immediately calculated from the first-order solution,

$$
\begin{equation*}
\zeta_{2}^{(1)}=\left[\frac{1}{g^{2}} \frac{\partial \Phi_{1}}{\partial t} \frac{\partial^{2} \Phi_{1}}{\partial t \partial z}-\frac{1}{2 g}\left(\nabla_{3} \Phi_{1}\right)^{2}\right]_{z=0} \tag{2.11}
\end{equation*}
$$

and the second part depends on the second-order potential $\Phi_{2}$

$$
\begin{equation*}
\zeta_{2}^{(2)}=\left[-\frac{1}{g} \frac{\partial \Phi_{2}}{\partial t}\right]_{z=0} \tag{2.12}
\end{equation*}
$$

whose solution is a major objective of this work.

## 3. Mild-slope equations

For convenience, we first cite the known result for the first order (Chamberlain \& Porter 1995).

### 3.1. First-order MSE

We assume the incident wave to be simple-harmonic in time at first order with the velocity potential

$$
\begin{equation*}
\Phi_{1}=\phi \mathrm{e}^{-\mathrm{i} \omega t}+* \tag{3.1}
\end{equation*}
$$

where * denotes the complex conjugate of the preceding term, and $\omega$ the wave frequency.

For vertical lateral boundaries, we take the first-order $\phi$ to be

$$
\begin{equation*}
\phi=-\frac{\mathrm{i} g \eta(x, y)}{\omega} \frac{\cosh k(z+h)}{\cosh k h} \tag{3.2}
\end{equation*}
$$

where $\eta(x, y)$ is the surface elevation for $\omega$ and $k$ is the wavenumber satisfying the dispersion relation

$$
\begin{equation*}
\omega^{2}=g k \tanh k h \tag{3.3}
\end{equation*}
$$

Since (3.2) representing only the propagating mode is exact for a horizontal seabed, it should be a good approximation for a mildly sloping bed.

By treating Laplace's equation as an ordinary differential equation in $z$ and applying Green's formula (as in Smith \& Sprinks 1975), we obtain the modified mild-slope equation for the first-order problem. The result by keeping all terms proportional to $\nabla h, \nabla^{2} h$ and $(\nabla h)^{2}$ is the modified mild-slope equation by Chamberlain \& Porter (1995),

$$
\begin{equation*}
\nabla \cdot\left(C C_{g} \nabla \eta\right)+\left[k^{2} C C_{g}+g U \nabla^{2} h+g V(\nabla h)^{2}\right] \eta=0, \tag{3.4}
\end{equation*}
$$

where

$$
\begin{align*}
U= & \frac{(\sinh 2 k h-2 k h \cosh 2 k h)}{4 \cosh ^{2}(k h)(2 k h+\sinh 2 k h)},  \tag{3.5}\\
V= & \frac{k\left[(2 k h)^{4}+4(2 k h)^{3} \sinh 2 k h-9 \sinh 2 k h \sinh 4 k h\right]}{12 \cosh ^{2}(k h)(2 k h+\sinh 2 k h)^{3}} \\
& +\frac{k\left[k h(k h+\sinh 2 k h)\left(\cosh ^{2} 2 k h-2 \cosh 2 k h+3\right)\right]}{\cosh ^{2}(k h)(2 k h+\sinh 2 k h)^{3}} \tag{3.6}
\end{align*}
$$

$C$ is the phase velocity

$$
\begin{equation*}
C=\frac{\omega}{k} \tag{3.7}
\end{equation*}
$$

and $C_{g}$ is the group velocity,

$$
\begin{equation*}
C_{g}=\frac{C}{2}\left(1+\frac{2 k h}{\sinh 2 k h}\right) \tag{3.8}
\end{equation*}
$$

Note that while no terms related to the gradients of depth are omitted, the assumed potential (3.2) is not exact and does not satisfy the bottom condition. The above approximation is therefore still restricted to mild slope.

Equation (3.4) with the appropriate radiation condition can be solved numerically by the hybrid element method of Chen \& Mei (1974) if the complex geometry is limited to a finite near field. Afterwards, the first-order free-surface elevation is given by

$$
\begin{equation*}
\zeta_{1}=\eta \mathrm{e}^{-\mathrm{i} \omega t}+* . \tag{3.9}
\end{equation*}
$$

With these results, the following part of the second-order free-surface displacement can be computed immediately,

$$
\begin{equation*}
\zeta_{2}^{(1)}=\eta_{2,0}^{(1)}+\left(\eta_{2,2}^{(1)} \mathrm{e}^{-\mathrm{i} 2 \omega t}+*\right) \tag{3.10}
\end{equation*}
$$

In particular, $\eta_{2,0}^{(1)}$ represents the mean sea-level set-up/setdown

$$
\begin{equation*}
\eta_{2,0}^{(1)}=\frac{\omega^{2}}{g}|\eta|^{2}-\frac{g}{\omega^{2}}|\nabla \eta|^{2}, \tag{3.11}
\end{equation*}
$$

and $\eta_{2,2}^{(1)}$ represents the second-harmonic amplitude due directly to quadratic interactions of the first-order motion

$$
\begin{equation*}
\eta_{2,2}^{(1)}=\frac{3 \omega^{2}}{2 g} \eta^{2}+\frac{g}{2 \omega^{2}}(\nabla \eta)^{2} \tag{3.12}
\end{equation*}
$$

Solution of the remaining second harmonic,

$$
\begin{equation*}
\zeta_{2}^{(2)}=\eta_{22}^{(2)} \mathrm{e}^{-2 \mathrm{i} \omega t}+* \tag{3.13}
\end{equation*}
$$

must await the solution of $\Phi_{2}$.

### 3.2. Second-order MSE

At second-order, the inhomogeneous free-surface boundary condition, (2.8), can be rewritten as follows

$$
\begin{equation*}
\frac{\partial \Phi_{2}}{\partial z}+\frac{1}{g} \frac{\partial^{2} \Phi_{2}}{\partial t^{2}}=F \mathrm{e}^{-2 \mathrm{i} \omega t}+* \tag{3.14}
\end{equation*}
$$

where

$$
\begin{equation*}
F=\widehat{\beta} \eta \eta+\bar{\beta} \nabla \eta \cdot \nabla \eta \tag{3.15}
\end{equation*}
$$

with

$$
\begin{equation*}
\widehat{\beta}=\frac{\mathrm{i} g k^{2}}{\omega}-\frac{3 \mathrm{i} \omega^{3}}{g}, \quad \bar{\beta}=-\frac{2 \mathrm{i} g}{\omega} . \tag{3.16}
\end{equation*}
$$

Expressing the second-order potential in the form

$$
\begin{equation*}
\Phi_{2}=\psi \mathrm{e}^{-2 \mathrm{i} \omega t}+* \tag{3.17}
\end{equation*}
$$

it is easy to see that $\psi$ must satisfy $\partial \psi / \partial n=0$ along lateral boundaries, (2.4) on the seabed and

$$
\begin{equation*}
\frac{\partial \psi}{\partial z}-\frac{4 \omega^{2}}{g} \psi=F, \quad z=0 \tag{3.18}
\end{equation*}
$$

on the free surface.
Similar to the first-order mild-slope approximation of Porter \& Staziker (1995), we express the solution for $\psi$ as the sum of all vertical eigenmodes in a sea of constant depth

$$
\begin{equation*}
\psi=-\frac{\mathrm{i} g}{2 \omega} \sum_{\ell=0}^{\infty} \xi_{\ell} \frac{\cos \kappa_{\ell}(z+h)}{\cos \kappa_{\ell} h} \tag{3.19}
\end{equation*}
$$

where $\kappa_{\ell}, \ell=1,2, \ldots$ are the real roots of the transcendental equation

$$
\begin{equation*}
-4 \omega^{2}=g \kappa_{\ell} \tan \kappa_{\ell} h, \quad\left(\ell-\frac{1}{2}\right) \pi \leqslant \kappa_{\ell} h \leqslant \ell \pi, \tag{3.20}
\end{equation*}
$$

and $\kappa_{0}=-\mathrm{i} \widehat{\kappa}_{0}$ is imaginary with $\widehat{\kappa}_{0}$ being the real root of the dispersion equation

$$
\begin{equation*}
4 \omega^{2}=g \widehat{\kappa}_{0} \tanh \widehat{\kappa}_{0} h \tag{3.21}
\end{equation*}
$$

Physically, the term with the subscript $m=0$ corresponds to the propagating mode while those with $m \geqslant 1$ correspond to the evanescent modes. Because of the inhomogeneous condition (3.18) on the free surface, all vertical modes are required, unlike the first-order potential in (3.2). Employing the procedure of weighted vertical averaging via Green's formula, as outlined in Appendix A, we obtain a matrix equation coupling all second-order modal amplitudes $\xi_{\ell}$ :

$$
\begin{equation*}
\sum_{\ell=0}^{\infty}\left\{\nabla \cdot\left(A_{m, \ell} \nabla \xi_{\ell}\right)+B_{m, \ell} \nabla h \cdot \nabla \xi_{\ell}+C_{m, \ell} \xi_{\ell}\right\}=-\mathrm{i} 2 \omega F, \quad m=0,1,2,3, \ldots \tag{3.22}
\end{equation*}
$$

The coefficient matrices defined below are functions of $(x, y)$,

$$
\begin{gather*}
A_{m, \ell}=\delta_{m \ell} A_{\ell, \ell}=\frac{g h}{2 \cos ^{2} \kappa_{\ell} h}\left(1+\frac{\sin 2 \kappa_{\ell} h}{2 \kappa_{\ell} h}\right) \delta_{m \ell},  \tag{3.23}\\
B_{m, \ell}= \begin{cases}0 & \text { for } m=\ell \\
g\left(U_{m, \ell}-U_{\ell, m}\right) & \text { for } m \neq \ell\end{cases} \tag{3.24}
\end{gather*}
$$



Figure 1. Near $\left(\Omega_{A}\right)$ and far ( $\Omega_{F}$ ) fields. Variable bathymetry and non-straight coastal boundaries are in $\left(\Omega_{A}\right)$ behind the semi-circle $r=a$.
with

$$
\begin{gather*}
U_{m, \ell}=\left\{\begin{array}{lc}
\frac{\sin 2 \kappa_{m} h-2 \kappa_{m} h \cos 2 \kappa_{m} h}{4 \cos ^{2}\left(\kappa_{m} h\right)\left(2 \kappa_{m} h+\sin 2 \kappa_{m} h\right)} & \text { if } m=\ell \\
-\frac{\kappa_{\ell}^{2}}{\cos \kappa_{m} h \cos \kappa_{\ell} h\left(\kappa_{\ell}^{2}-\kappa_{m}^{2}\right)}, & \text { if } m \neq \ell
\end{array}\right.  \tag{3.25}\\
C_{m, \ell}=-\kappa_{m}^{2} A_{m, \ell}+g U_{m, \ell} \nabla^{2} h+g V_{m, \ell}(\nabla h)^{2} \tag{3.26}
\end{gather*}
$$

with

$$
\begin{align*}
V_{m, m}= & \frac{\kappa_{m}\left[-\left(2 \kappa_{m} h\right)^{4}-4\left(2 \kappa_{m} h\right)^{3} \sin 2 \kappa_{m} h-9 \sin \left(2 \kappa_{m} h\right) \sin 4 \kappa_{m} h\right]}{12 \cos ^{2}\left(\kappa_{m} h\right)\left(2 \kappa_{m} h+\sin 2 \kappa_{m} h\right)^{3}} \\
& +\frac{\kappa_{m}\left[\kappa_{m} h\left(\kappa_{m} h+\sin 2 \kappa_{m} h\right)\left(\cos ^{2} 2 \kappa_{m} h-2 \cos 2 \kappa_{m} h+3\right)\right]}{\cos ^{2}\left(\kappa_{m} h\right)\left(2 \kappa_{m} h+\sin 2 \kappa_{m} h\right)^{3}} \tag{3.27}
\end{align*}
$$

for $m=\ell$, and

$$
\begin{equation*}
V_{m, \ell}=\frac{-2 \kappa_{\ell} \sec \kappa_{m} h \sec \kappa_{\ell} h}{\left(2 \kappa_{\ell} h+\sin 2 \kappa_{\ell} h\right)} \frac{\left[4 \kappa_{\ell}^{2} \kappa_{m}^{2}+\left(\kappa_{\ell}^{4}-\kappa_{m}^{4}\right) \sin ^{2} \kappa_{\ell} h\right]}{\left(\kappa_{\ell}^{2}-\kappa_{m}^{2}\right)^{2}} \tag{3.28}
\end{equation*}
$$

for $m \neq \ell$. These coefficients are equivalent to those in the modified system of mildslope equations of Porter \& Staziker (1995) for linearized waves and $F=0$, after changing $2 \omega$ to $\omega$. We have now reduced the second-order task to the solution of coupled two-dimensional elliptic problems, subject to certain boundary conditions.

In the region of constant depth, terms multiplied by $B_{m \ell}$ vanish, while $C_{m, \ell}$ is also a diagonal matrix. Equations (3.22) are no longer coupled and reduce to

$$
\begin{equation*}
\nabla^{2} \xi_{\ell}-\kappa_{\ell}^{2} \xi_{\ell}=-\mathrm{i} \frac{2 \omega}{A_{\ell, \ell}} F, \quad \ell=0,1,2, \ldots \tag{3.29}
\end{equation*}
$$

For coastal waters where only the complexities of local topography and coastline are of major concern, we can simplify the geometry in areas far from the local region. With reference to figure 1 , let us divide the horizontal fluid domain into two regions : the near field $\Omega_{A}$ in which the bathymetry and coastal boundary are complex, and the far field $\Omega_{F}$ where the depth is constant and the coastline straight. The two fields are separated by a semi-circle $\partial A$ of radius $r=a$. For the first-order problem, the far-field solution in $\Omega_{F}$ can be represented analytically as an eigenfunction expansion. In the near field $\Omega_{A}$, discrete solutions will be sought via finite elements. The unknown nodal coefficients in $\Omega_{A}$ and the expansion coefficients in $\Omega_{F}$ will be found together
by Galerkin's method, subject to the continuity of $\eta$ and $\partial \eta / \partial r$ at $r=a$. This is the idea of the hybrid-element method developed for linearized wave problems by Chen \& Mei (1974) and Bai \& Yeung (1974) for two dimensions and Yue, Chen \& Mei (1978) for three dimensions. The method will be extended to the second-order problem here with the requirement of continuity of $\xi_{\ell}$ and $\partial \xi_{\ell} / \partial r$ at $r=a$.

## 4. Analytical representations in the far field

As a part of the hybrid-element analysis, let us first present the analytical representations in the far field $\Omega_{F}$ where the depth is constant and the coastline is along the $x$-axis.

### 4.1. First order

In the far field, the first-order wave consists of the incident, reflected and scattered waves given together by

$$
\begin{equation*}
(\eta)_{\Omega_{F}}=\eta^{(T)}+\eta^{(S)} \tag{4.1}
\end{equation*}
$$

where

$$
\begin{equation*}
\eta^{(T)}=\frac{1}{2} A \exp \left(\mathrm{i} k r \cos \left(\theta-\theta_{I}\right)\right)+\frac{1}{2} A \exp \left(\mathrm{i} k r \cos \left(\theta+\theta_{I}\right)\right) \tag{4.2}
\end{equation*}
$$

is the sum of incident and reflected waves in the presence of the straight coast, with $\theta_{I}$ being the incidence angle. The first-order scattered wave $\eta^{(S)}$ is formally

$$
\begin{equation*}
\eta^{(S)}=\sum_{m=0}^{\infty} \epsilon_{m} \mathrm{i}^{m} \alpha_{m} H_{m}(k r) \cos m \theta \tag{4.3}
\end{equation*}
$$

where $\epsilon_{m}$ is the Jacobi symbol, being 1 for $m=0$ and 2 for $m=1,2,3, \ldots$, $H_{m}(k r) \equiv H_{m}^{(1)}(k r)$ is a Hankel function of the first kind, and the coefficients $\alpha_{m}, m=0,1,2, \ldots$, are as yet unknown.

In the near field of complex bathymetry and coastline, discrete finite elements are used. The coupled problems of near and far fields are solved numerically via a variational principle as in Chen \& Mei (1974) as described by Houston (1981). Afterwards the nodal values of $\eta$ in the near field and the coefficients $\alpha_{m}$ of the far field are found.

### 4.2. Second order

### 4.2.1. The forcing function $F$

Referring to (3.15), $F$ contains quadratic products of the first-order waves and can be decomposed as follows

$$
\begin{equation*}
F=\mathscr{P}+\mathscr{Q} \tag{4.4}
\end{equation*}
$$

where $\mathscr{P}$ denotes the part associated with the self- and cross-interactions of progressive waves (incident and reflected waves),

$$
\begin{equation*}
\mathscr{P}=\widehat{\beta} \eta^{(T)} \eta^{(T)}+\bar{\beta} \nabla \eta^{(T)} \cdot \nabla \eta^{(T)} \tag{4.5}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathscr{2}=\widehat{\beta}\left[2 \eta^{(T)}+\eta^{(S)}\right] \eta^{(S)}+\bar{\beta}\left[2 \nabla \eta^{(T)}+\nabla \eta^{(S)}\right] \cdot \nabla \eta^{(S)} \tag{4.6}
\end{equation*}
$$

is the part due to cross-interactions between the plane progressive waves and the scattered waves, and the self-interaction of the scattered waves.

By straightforward calculations, we obtain

$$
\begin{align*}
\mathscr{P}=\mathscr{M}+\mathscr{N}= & \frac{1}{4} A^{2}\left(\widehat{\beta}-\bar{\beta} k^{2}\right)\left[\exp \left(\mathrm{i} 2 k r \cos \left(\theta-\theta_{I}\right)\right)+\exp \left(\mathrm{i} 2 k r \cos \left(\theta+\theta_{I}\right)\right)\right] \\
& +\frac{1}{2} A^{2}\left[\widehat{\beta}-\bar{\beta} k^{2} \cos \left(2 \theta_{I}\right)\right] \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta\right) \tag{4.7}
\end{align*}
$$

Clearly, the first line,

$$
\begin{equation*}
\mathscr{M}=\frac{1}{4} A^{2}\left(\widehat{\beta}-\bar{\beta} k^{2}\right)\left[\exp \left(\mathrm{i} 2 k r \cos \left(\theta-\theta_{I}\right)\right)+\exp \left(\mathrm{i} 2 k r \cos \left(\theta+\theta_{I}\right)\right)\right] \tag{4.8}
\end{equation*}
$$

represents a pair of obliquely incident and reflected plane waves of wavenumber $2 k$, and the second line,

$$
\begin{equation*}
\mathscr{N}=\frac{1}{2} A^{2}\left[\widehat{\beta}-\bar{\beta} k^{2} \cos \left(2 \theta_{I}\right)\right] \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta\right) \tag{4.9}
\end{equation*}
$$

is a plane incident wave of wavenumber $2 k \cos \theta_{I}$ propagating along the coastline.
To calculate the forcing function $\mathscr{Q}$, we first represent the first-order progressive wave as a partial wave expansion

$$
\begin{equation*}
\eta^{(T)}=\sum_{m=-\infty}^{\infty} T_{m} \mathrm{e}^{(\mathrm{i} m \theta)} \text { with } T_{m}(r)=A \mathrm{i}^{m} J_{m}(k r) \cos m \theta_{I} \tag{4.10}
\end{equation*}
$$

and the first-order scattered waves as

$$
\begin{equation*}
\eta^{(S)}=\sum_{m=-\infty}^{\infty} S_{m} \mathrm{e}^{\mathrm{i} m \theta} \quad \text { where } S_{m}(r)=\mathrm{i}^{m} \alpha_{m} H_{m}(k r) \tag{4.11}
\end{equation*}
$$

where the coefficients $\alpha_{m}$ are to be solved by the hybrid-element method. Because of the no flux condition on the straight coast,

$$
\begin{equation*}
\frac{1}{r} \frac{\partial\left(\eta^{(T)}, \eta^{(S)}\right)}{\partial \theta}=0, \quad \theta=0, \pi \tag{4.12}
\end{equation*}
$$

$\alpha_{m}, T_{m}, S_{m}$ must be even in $m$.
We now calculate 2 according to (4.6). A typical cross-product in (4.6) is of the form

$$
\begin{equation*}
\eta^{(T)} \eta^{(S)}=\sum_{m=-\infty}^{\infty} T_{m} \mathrm{e}^{\mathrm{i} m \theta} \sum_{n=-\infty}^{\infty} S_{n} \mathrm{e}^{\mathrm{i} n \theta}=\sum_{m=-\infty}^{\infty}\left[\sum_{n=-\infty}^{\infty} T_{m-n} S_{n}\right] \mathrm{e}^{\mathrm{i} m \theta} \tag{4.13}
\end{equation*}
$$

It follows that 2 can be expressed as a Fourier series

$$
\begin{equation*}
\mathscr{Q}=\sum_{m=-\infty}^{\infty} \mathscr{Z}_{m} \mathrm{e}^{\mathrm{i} m \theta}=\sum_{m=0}^{\infty} \epsilon_{m} \mathscr{Q}_{m}(r) \cos m \theta \tag{4.14}
\end{equation*}
$$

with

$$
\begin{equation*}
\mathscr{Q}_{m}=\sum_{n=-\infty}^{\infty}\left\{\left[\widehat{\beta}-\frac{n(m-n)}{r^{2}} \bar{\beta}\right] S_{n}\left(2 T_{m-n}+S_{m-n}\right)+\bar{\beta} \frac{\partial S_{n}}{\partial r}\left(2 \frac{\partial T_{m-n}}{\partial r}+\frac{\partial S_{m-n}}{\partial r}\right)\right\} \tag{4.15}
\end{equation*}
$$

4.2.2. Decomposition of the far-field solution

In accordance with the form of the forcing term, the second-order response $\left(\xi_{\ell}\right)_{\Omega_{F}}$ defined by (3.19) can be separated into three parts,

$$
\begin{equation*}
\left(\xi_{\ell}\right)_{\Omega_{F}}=\xi_{\ell}^{P}+\xi_{\ell}^{Q}+\xi_{\ell}^{H} \tag{4.16}
\end{equation*}
$$

The first part, $\xi_{\ell}^{P}$, is the direct response to $\mathscr{P}$. Thus, it satisfies the following inhomogeneous equation

$$
\begin{equation*}
\left(\nabla^{2}-\kappa_{\ell}^{2}\right) \xi_{\ell}^{P}=-\mathrm{i} \frac{2 \omega \mathscr{P}}{A_{\ell, \ell}}, \quad \ell=0,1,2,3, \ldots \tag{4.17}
\end{equation*}
$$

The above decomposition is not unique. It will be shown in the next subsection that the solution is just the sum of three Stokes' waves.

The second part, $\xi_{\ell}^{Q}$, is the response to forcing by $\mathscr{Q}$ and is required to satisfy the inhomogeneous equation

$$
\begin{equation*}
\left(\nabla^{2}-\kappa_{\ell}^{2}\right) \xi_{\ell}^{Q}=-i \frac{2 \omega \mathscr{Q}}{A_{\ell, \ell}}, \quad \ell=0,1,2,3, \ldots, \tag{4.18}
\end{equation*}
$$

the no-flux condition along the straight coast

$$
\begin{equation*}
\frac{\partial \xi_{\ell}^{Q}}{\partial \theta}=0, \quad r>a ; \quad \theta=0, \quad \theta=\pi, \quad \ell=0,1,2,3, \ldots \tag{4.19}
\end{equation*}
$$

and along the semi-circle $r=a$

$$
\begin{equation*}
\frac{\partial \xi_{\ell}^{Q}}{\partial r}=0, \quad r=a, \quad 0 \leqslant \theta \leqslant \pi, \quad \ell=0,1,2,3, \ldots \tag{4.20}
\end{equation*}
$$

This condition is chosen for convenience. For the evanescent modes, $\ell=1,2,3, \ldots$, $\xi_{\ell}^{Q}$ diminishes to zero at large enough $r$. For the propagating mode, $\ell=0, \xi_{0}^{Q}$ must satisfy the weak (integral) radiation condition at infinity owing to the slow attenuation of $Q$. We shall call both $\xi_{\ell}^{P}$ and $\xi_{\ell}^{Q}$ the forced waves.

Finally, $\xi_{\ell}^{H}$ is the solution to the homogeneous Helmholtz equation

$$
\begin{equation*}
\left(\nabla^{2}-\kappa_{\ell}^{2}\right) \xi_{\ell}^{H}=0, \quad \ell=0,1,2,3, \ldots \tag{4.21}
\end{equation*}
$$

and the no-flux condition along the straight coast,

$$
\begin{equation*}
\frac{\partial \xi_{\ell}^{H}}{\partial \theta}=0, \quad r>a, \quad \theta=0 \quad \theta=\pi, \quad \ell=0,1,2,3, \ldots \tag{4.22}
\end{equation*}
$$

In addition, $\xi_{\ell}^{H}$ must satisfy the usual (strong) radiation condition at infinity. To be referred to as the free wave, the formal solution $\xi_{\ell}^{H}$ is immediate

$$
\begin{equation*}
\xi_{\ell}^{H}=\sum_{m=0}^{\infty} \epsilon_{m} \widehat{\alpha}_{\ell, m} K_{m}\left(\kappa_{\ell} r\right) \cos m \theta, \quad \ell=0,1,2,3, \ldots, \tag{4.23}
\end{equation*}
$$

where $K_{m}$ is the modified Bessel function of the second kind of order $m$. For $\ell=0, \kappa_{0}=-\mathrm{i} \widehat{\kappa}_{0}$ is imaginary and $K_{m}\left(\kappa_{0} r\right)$ is proportional to $H_{m}^{(1)}\left(\widehat{\kappa}_{0} r\right)$. The unknown coefficients $\widehat{\alpha}_{\ell, m}$ will be found jointly with the discrete solution in the near field by the hybrid-element analysis to be described shortly, which requires the continuity of potentials and their radial derivatives along $r=a$.

The solutions for $\xi_{\ell}^{P}$ and $\xi_{\ell}^{Q}$ can be obtained explicitly, as shown below.

### 4.2.3. Response $\xi_{\ell}^{P}$ to forcing $\mathscr{P}$ by progressive waves

It is easy to derive from (4.17) that

$$
\begin{equation*}
\xi_{\ell}^{P}=\xi_{\ell}^{I}+\xi_{\ell}^{R}+\xi_{\ell}^{I R}, \quad \ell=0,1,2,3, \ldots \tag{4.24}
\end{equation*}
$$

where

$$
\begin{align*}
\xi_{\ell}^{I} & =L_{\ell}^{M} \exp \left(\mathrm{i} 2 k r \cos \left(\theta-\theta_{I}\right)\right), \quad \xi_{\ell}^{R}=L_{\ell}^{M} \exp \left(\mathrm{i} 2 k r \cos \left(\theta+\theta_{I}\right)\right) \\
\xi_{\ell}^{I R} & =L_{\ell}^{N} \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta\right) \tag{4.25}
\end{align*}
$$

with

$$
\begin{equation*}
L_{\ell}^{M}=\frac{2 \mathrm{i} \omega}{A_{\ell, \ell}} \frac{A^{2}}{4} \frac{\left(\widehat{\beta}-\bar{\beta} k^{2}\right)}{4 k^{2}+\kappa_{\ell}^{2}}, \quad L_{\ell}^{N}=\frac{2 \mathrm{i} \omega}{A_{\ell, \ell}} \frac{A^{2}}{2} \frac{\left[\widehat{\beta}-\bar{\beta} k^{2} \cos \left(2 \theta_{I}\right)\right]}{\left(2 k \cos \theta_{I}\right)^{2}+\kappa_{\ell}^{2}} \tag{4.26}
\end{equation*}
$$

Substituting (4.25) into (3.19), we obtain the corresponding second-order potentials $\Phi_{2}^{I}, \Phi_{2}^{R}$ and $\Phi_{2}^{I R}$,

$$
\begin{align*}
\binom{\Phi_{2}^{I}}{\Phi_{2}^{R}}= & {\left[\sum_{\ell=0}^{\infty} \frac{g}{A_{\ell, \ell}\left(4 k^{2}+\kappa_{\ell}^{2}\right)} \frac{\cos \kappa_{\ell}(z+h)}{\cos \kappa_{\ell} h}\right] } \\
& \times\left(\widehat{\beta}-\bar{\beta} k^{2}\right) \frac{A^{2}}{4} \exp \left(\mathrm{i} 2 k r \cos \left(\theta \mp \theta_{I}\right)-\mathrm{i} 2 \omega t\right)+* \tag{4.27}
\end{align*}
$$

and

$$
\begin{align*}
\Phi_{2}^{I R}= & {\left[\sum_{\ell=0}^{\infty} \frac{g}{A_{\ell, \ell}\left(\left(2 k \cos \theta_{I}\right)^{2}+\kappa_{\ell}^{2}\right)} \frac{\cos \kappa_{\ell}(z+h)}{\cos \kappa_{\ell} h}\right] } \\
& \times\left(\widehat{\beta}-\bar{\beta} k^{2} \cos 2 \theta_{I}\right) \frac{A^{2}}{2} \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta-\mathrm{i} 2 \omega t\right)+* \tag{4.28}
\end{align*}
$$

For normal incidence $\theta_{I}=3 \pi / 2$, the second-harmonic potential $\Phi_{2}^{I R}$ becomes independent of $(x, y, z)$ (or $r, \theta, z$ ); the associated second harmonic pressure persists for all depth down to the seabed and can induce microseisms, as shown first by Longuet-Higgins (1950) for a pure standing wave in deep water on a straight coast.

In Appendix B , it is shown that $\Phi_{2}^{I}, \Phi_{2}^{R}$ and $\Phi_{2}^{I R}$ are just the series expansions of the second-order part of the classical Stokes waves. (Strictly speaking, the classical Stokes wave is a single wave train. The three-term result here is the second-order part of an obliquely incident and reflected wave system.) While the Stokes forms are more compact, the series forms here are more convenient for later computations.

### 4.2.4. Response $\xi_{\ell}^{Q}$ to forcing by $Q$

To solve the inhomogeneous problem for each $\xi_{\ell}^{Q}$, we shall employ two-dimensional Green's functions, $G_{\ell}\left(r, \theta ; r_{0}, \theta_{0}\right)$, described in Appendix C. By Green's theorem, the boundary conditions and the reciprocity of Green's functions, we find at any field point ( $r, \theta$ ),

$$
\begin{align*}
\xi_{\ell}^{Q}(r, \theta)= & \int_{a}^{\infty} r_{0} \mathrm{~d} r_{0} \int_{0}^{\pi} \mathrm{d} \theta_{0}\left[-\mathrm{i} \frac{2 \omega \mathscr{2}\left(r_{0}, \theta_{0}\right)}{A_{\ell, \ell}}\right] \sum_{m=0}^{\infty} \frac{\epsilon_{m}}{\pi} \cos m \theta \cos m \theta_{0} K_{m}\left(\kappa_{\ell} r_{>}\right) \\
& \times\left[\frac{I_{m}^{\prime}\left(\kappa_{\ell} a\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r_{<}\right)-I_{m}\left(\kappa_{\ell} r_{<}\right)\right], \quad \ell=0,1,2,3, \ldots \tag{4.29}
\end{align*}
$$

where

$$
\begin{equation*}
r_{<}=\min \left(r, r_{0}\right), \quad r_{>}=\max \left(r, r_{0}\right) \tag{4.30}
\end{equation*}
$$

The series above is the Green function $G_{\ell}\left(r, \theta ; r_{0}, \theta_{0}\right)$. Use has been made of the facts that (i) for $\ell=1,2,3, \ldots$, the evanescent modes die out exponentially at infinity, and (ii) for $\ell=0$, the propagating mode satisfies the weak radiation condition, i.e. the integral below vanishes in the limit of large $r$,

$$
\begin{equation*}
\int_{\partial \Omega_{\infty}}\left(\xi_{0} \frac{\partial G_{0}}{\partial n}-G_{0} \frac{\partial \xi_{0}}{\partial n}\right) \mathrm{d} S=O\left(\frac{1}{\sqrt{r}}\right) \rightarrow 0, \quad r \rightarrow \infty \tag{4.31}
\end{equation*}
$$

Verification of this result is similar to that given in Mei (1989, p. 664). Using the Fourier expansion (4.14) for $\mathscr{2}(r, \theta)$, the orthogonality of cosines, and the Wronskian
identity of Bessel functions, we obtain the value on the circle $r=a$,

$$
\begin{equation*}
\xi_{\ell}^{Q}(a, \theta)=\sum_{m=0}^{\infty} \frac{\epsilon_{m}}{\kappa_{\ell} a} \cos (m \theta) \int_{a}^{\infty} r_{0} \mathrm{~d} r_{0}\left[-\mathrm{i} \frac{2 \omega \mathscr{2}_{m}\left(r_{0}\right)}{A_{\ell, \ell}}\right] \frac{K_{m}\left(\kappa_{\ell} r_{0}\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)}, \quad \ell=0,1,2,3, \ldots \tag{4.32}
\end{equation*}
$$

which is needed to calculate the finite-element solution inside $\Omega_{A}$. The numerical evaluation of the integral above can be expedited by the scheme of Chau \& Eatock Taylor (1992).

In summary, the second-order solution in the far-field is known analytically except for the unknown coefficients $\widehat{\alpha}_{\ell, m}$ of the second-order free wave.

## 5. Hybrid-element solution

In the near field $\Omega_{A}$ of complex geometry, the amplitudes $\xi_{\ell}(x, y)$ (without decomposition) satisfy (3.22) where the forcing function $F(x, y)$ is known at every nodal point from the numerical solution of the first-order problem. The no-flux condition must be imposed on the lateral boundary $\partial B$ which is the union of the complex coastline (peninsula etc.),

$$
\begin{equation*}
\left(\frac{\partial \xi_{\ell}}{\partial n}\right)_{\Omega_{A}}=0 \quad \text { on } \partial B, \quad \ell=0,1,2,3, \ldots \tag{5.1}
\end{equation*}
$$

In addition we require the continuity of pressure along and flux across the semi-circle $\partial A:(r=a, 0 \leqslant \theta \leqslant \pi)$

$$
\begin{array}{cc}
\left(\xi_{\ell}\right)_{\Omega_{A}}=\xi_{\ell}^{P}+\xi_{\ell}^{Q}+\xi_{\ell}^{H}, & r=a, \\
\left(\frac{\partial \xi_{\ell}}{\partial r}\right)_{\Omega_{A}}=\frac{\partial \xi_{\ell}^{P}}{\partial r}+\frac{\partial \xi_{\ell}^{H}}{\partial r}, \quad r=a, 1,2,3, \ldots \tag{5.3}
\end{array}
$$

Use has been made of (4.20). These conditions are imposed on the nodal points along the semicircle. Recall that the far-field solutions $\xi_{\ell}^{P}, \xi_{\ell}^{Q}$ and $\xi_{\ell}^{H}$ are given by (4.24), (4.32) and (4.23), the last of which still contains the unknown coefficients $\widehat{\alpha}_{\ell, m}$.

### 5.1. Finite-element analysis

Equation (3.22) is first truncated after a finite number of $N_{\xi}+1$ modes; that is, $(\ell, m)=0,1, \ldots, N_{\xi}$. Afterwards we applied the method of weighted residuals and form the integrals for the truncated equation

$$
\begin{equation*}
\iint_{\Omega_{A}}\left\{-\sum_{\ell=0}^{N_{\xi}}\left[\nabla \cdot\left(A_{m, \ell} \nabla \xi_{\ell}\right)+B_{m, \ell} \nabla h \cdot \nabla \xi_{\ell}+C_{m, \ell} \xi_{\ell}\right]=\mathrm{i} 2 \omega F\right\} \mathscr{W} \mathrm{d} \Omega \tag{5.4}
\end{equation*}
$$

where $\mathscr{W}$ is a weighting function to be chosen shortly. By partial integration, and the use of boundary conditions (5.1) and (5.3), the preceding equation becomes

$$
\begin{gather*}
\iint_{\Omega_{A}}\left[\sum_{\ell=0}^{N_{\xi}}\left(A_{m, \ell} \nabla \xi_{\ell} \cdot \nabla \mathscr{W}-B_{m, \ell} \mathscr{W} \nabla h \cdot \nabla \xi_{\ell}-C_{m, \ell} \xi_{\ell} \mathscr{W}\right)\right] \mathrm{d} \Omega \\
=\int_{0}^{\pi} \mathscr{W} A_{m, m}\left(\frac{\partial \xi_{m}^{P}}{\partial r}+\frac{\partial \xi_{m}^{H}}{\partial r}\right) a \mathrm{~d} \theta+\iint_{\Omega_{A}} \mathrm{i} 2 \omega F \mathscr{W} \mathrm{~d} \Omega, \quad m=0,1,2,3, \ldots, N_{\xi} \tag{5.5}
\end{gather*}
$$

Use has been made of (3.23) where $A_{\ell, m}$ is diagonal.
Now the domain $\Omega_{A}$ is discretized to an assemblage of three-node triangular elements $\left(\triangle_{e}\right)$ with a total of $N_{E}$ elements and $N_{P}$ nodes, among which $N_{B}$ boundary
nodes are designated to be at the end of the sequence, and are labelled as nodes $N_{P}-N_{B}+1$ to $N_{P}$. The unknown variable $\xi_{\ell}$ is replaced by

$$
\begin{equation*}
\xi_{\ell}=\sum_{n=1}^{N_{P}} \xi_{\ell, n} \mathscr{W}_{n}(x, y), \quad \ell=0,1,2,3, \ldots \tag{5.6}
\end{equation*}
$$

where $\xi_{\ell, n}$ denotes the $n$th unknown nodal coefficient of $\xi_{\ell}$ and $\mathscr{W}_{n}$ the global weighting function. In addition, we choose $\mathscr{W}_{\tau}$ with $\tau=1,2,3, \ldots, N_{p}$ to be the weighting function $\mathscr{W}$ in (5.5). Accordingly, (5.5) is approximated as

$$
\begin{gather*}
\sum_{\ell=0}^{N_{\xi}} \sum_{n=1}^{N_{P}} \xi_{\ell, n}\left\{\sum_{\Delta_{e}=1}^{N_{E}} \iint_{\Delta_{e}}\left[\left(A_{m, \ell} \nabla \mathscr{W}_{n} \cdot \nabla \mathscr{W}_{\tau}-B_{m, \ell} \mathscr{W}_{\tau} \nabla h \cdot \nabla \mathscr{W}_{n}-C_{m, \ell} \mathscr{W}_{n} \mathscr{W}_{\tau}\right)\right] \mathrm{d} \Omega\right\} \\
-\int_{0}^{\pi} \mathscr{W}_{\tau} A_{m, m} \frac{\partial \xi_{m}^{H}}{\partial r} a \mathrm{~d} \theta=\int_{0}^{\pi} \mathscr{W}_{\tau} A_{m, m} \frac{\partial \xi_{m}^{P}}{\partial r} a \mathrm{~d} \theta+\sum_{\Omega_{e}=1}^{N_{E}} \iint_{\Omega_{e}} \mathrm{i} 2 \omega F \mathscr{W}_{\tau} \mathrm{d} \Omega \\
m=0,1,2, \ldots, N_{\xi}, \quad \tau=1,2, \ldots, N_{p} \tag{5.7}
\end{gather*}
$$

Furthermore, the Fourier series expansion in the free wave $\xi_{\ell}^{H}$, (4.23), is truncated after $N_{\widehat{\alpha}}+1$ terms and also the dummy index $m$ for the Fourier series summation is changed to $p$. The preceding equation becomes

$$
\begin{align*}
& \sum_{\ell=0}^{N_{\xi}} \sum_{n=1}^{N_{P}} \xi_{\ell, n}\left\{\sum_{\Delta_{e}=1}^{N_{E}} \iint_{\Delta_{e}}\left(A_{m, \ell} \nabla \mathscr{W}_{n} \cdot \nabla \mathscr{W}_{\tau}-B_{m, \ell} \mathscr{W}_{\tau} \nabla h \cdot \nabla \mathscr{W}_{n}-C_{m, \ell} \mathscr{W}_{n} \mathscr{W}_{\tau}\right) \mathrm{d} \Omega\right\} \\
& \quad-\sum_{p=0}^{N_{\widehat{\alpha}}} \epsilon_{p} \widehat{\alpha}_{m, p}\left\{\left[a \kappa_{m} A_{m, m} K_{p}^{\prime}\left(\kappa_{m} a\right)\right] \int_{0}^{\pi} \mathscr{W}_{\tau} \cos p \theta \mathrm{~d} \theta\right\} \\
& = \\
& \quad \mathrm{i} 2 k a A_{m, m} L_{m}^{M} \int_{o}^{\pi} \mathscr{W}_{\tau}\left[\cos \left(\theta-\theta_{I}\right) \exp \left(\mathrm{i} 2 k a \cos \left(\theta-\theta_{I}\right)\right)+\cos \left(\theta+\theta_{I}\right)\right. \\
& \left.\quad \times \exp \left(\mathrm{i} 2 k a \cos \left(\theta+\theta_{I}\right)\right)\right] \mathrm{d} \theta+\mathrm{i} 2 k a A_{m, m} \cos \theta_{I} L_{m}^{N} \int_{0}^{\pi} \mathscr{W}_{\tau} \exp \left(\mathrm{i} 2 k a \cos \theta_{I} \cos \theta\right) \cos \theta \mathrm{d} \theta  \tag{5.8}\\
& \quad+\sum_{\Delta_{e}=1}^{N_{E}} \mathrm{i} 2 \omega \iint_{\Delta_{e}} F \mathscr{W}_{\tau} \mathrm{d} \Omega, \quad m=0,1,2, \ldots, N_{\xi}, \quad \tau=1,2, \ldots, N_{p}
\end{align*}
$$

after using the expression of $\xi_{\ell}^{P}$ (cf. (4.25)).
Let us introduce the following abbreviations,

$$
\begin{align*}
& \mathscr{K}_{\tau, n}^{(m, \ell)}= \sum_{\Delta_{e}=1}^{N_{E}} \iint_{\Delta_{e}}\left(A_{m, \ell} \nabla \mathscr{W}_{n} \cdot \nabla \mathscr{W}_{\tau}-B_{m, \ell} \mathscr{W}_{\tau} \nabla h \cdot \nabla \mathscr{W}_{n}-C_{m, \ell} \mathscr{W}_{n} \mathscr{W}_{\tau}\right) \mathrm{d} \Omega  \tag{5.9}\\
& \widehat{\mathscr{K}}_{\tau, p}^{(m, m)}=-\epsilon_{p} a \kappa_{m} A_{m, m} K_{p}^{\prime}\left(\kappa_{m} a\right) \int_{0}^{\pi} \mathscr{W}_{\tau} \cos p \theta \mathrm{~d} \theta  \tag{5.10}\\
& \mathscr{Y}_{\tau}^{(m)}= \mathrm{i} 2 k a A_{m, m} L_{m}^{M} \int_{0}^{\pi} \mathscr{W}_{\tau}\left[\cos \left(\theta-\theta_{I}\right) \exp \left(\mathrm{i} 2 k a \cos \left(\theta-\theta_{I}\right)\right)\right. \\
&\left.+\cos \left(\theta+\theta_{I}\right) \exp \left(\mathrm{i} 2 k a \cos \left(\theta+\theta_{I}\right)\right)\right] \mathrm{d} \theta+\mathrm{i} 2 k a A_{m, m} \cos \theta_{I} L_{m}^{N} \\
& \times \int_{0}^{\pi} \mathscr{W}_{\tau} \exp \left(\mathrm{i} 2 k a \cos \theta_{I} \cos \theta\right) \cos \theta \mathrm{d} \theta+\sum_{\Delta_{e}=1}^{N_{E}} \mathrm{i} 2 \omega \iint_{\Delta_{e}} F \mathscr{W}_{\tau} \mathrm{d} \Omega \tag{5.11}
\end{align*}
$$

where all the integrals above are evaluated numerically. Equation (5.8) can then be expressed more compactly as

$$
\begin{equation*}
\sum_{\ell=0}^{N_{\xi}} \sum_{n=1}^{N_{P}} \xi_{\ell, n} \mathscr{K}_{\tau, n}^{(m, \ell)}+\sum_{p=0}^{N_{\widetilde{\alpha}}} \widehat{\alpha}_{m, p} \widehat{\mathcal{K}}_{\tau, p}^{(m, m)}=\mathscr{Y}_{\tau}^{(m)}, \quad m=0,1,2, \ldots, N_{\xi}, \quad \tau=1,2, \ldots, N_{P} \tag{5.12}
\end{equation*}
$$

Since $m=0,1, \ldots, N_{\xi}$ and $\tau=1,2, \ldots, N_{P}$, these constitute $\left(N_{\xi}+1\right) \times N_{P}$ equati ons. There are, however, the same number of unknown nodal coefficients plus ( $N_{\xi}+$ 1) $\times\left(N_{\widehat{\alpha}}+1\right)$ unknown free-wave coefficients $\widehat{\alpha}_{\ell, m}$. The remaining equations are found from the matching condition (5.2) after using (5.6) and (4.23)

$$
\begin{equation*}
-\sum_{n=1}^{N_{P}} \xi_{\ell, n} \mathscr{W}_{n}+\sum_{p=0}^{N_{\widehat{\alpha}}} \epsilon_{p} \widehat{\alpha}_{\ell, p} K_{p}\left(\kappa_{\ell} a\right) \cos p \theta=-\xi_{\ell}^{P}-\xi_{\ell}^{Q}, \quad \ell=0,1,2, \ldots, N_{\xi} \tag{5.13}
\end{equation*}
$$

Let us multiply (5.13) by $\left[\epsilon_{q} A_{\ell, \ell} \kappa_{\ell} a K_{q}^{\prime}\left(\kappa_{\ell} a\right) \cos q \theta\right]$ with $q=0,1,2, \ldots, N_{\widehat{\alpha}}$ and integrate over $(0 \leqslant \theta \leqslant \pi)$. Since

$$
\begin{equation*}
\int_{0}^{\pi} \cos p \theta \cos q \theta \mathrm{~d} \theta=0 \quad \text { when } p \neq q \tag{5.14}
\end{equation*}
$$

we obtain
$\left[\sum_{n=1}^{N_{P}} \check{\mathscr{K}}_{q, n}^{(\ell, \ell)} \xi_{\ell, n}\right]+\widetilde{\mathscr{K}}_{q, q}^{(\ell, \ell)} \widehat{\alpha}_{\ell, q}=\widetilde{\mathscr{Y}}_{q}^{(\ell)}, \quad \ell=0,1,2, \ldots, N_{\xi}, \quad q=0,1,2, \ldots, N_{\widehat{\alpha}}$,
where

$$
\begin{gathered}
\check{\mathscr{K}}_{q, h}^{(\ell, \ell)}=-\epsilon_{q} A_{\ell, \ell} \kappa_{\ell} a K_{q}^{\prime}\left(\kappa_{\ell} a\right) \int_{0}^{\pi} \mathscr{W}_{n} \cos q \theta \mathrm{~d} \theta, \\
\widetilde{\mathscr{K}}_{q, q}^{(\ell, \ell)}=\epsilon_{q}^{2} A_{\ell, \ell} \kappa_{\ell} a K_{q}^{\prime}\left(\kappa_{\ell} a\right) K_{q}\left(\kappa_{\ell} a\right) \int_{0}^{\pi} \cos ^{2} q \theta \mathrm{~d} \theta=\epsilon_{q} A_{\ell, \ell} \kappa_{\ell} a K_{q}^{\prime}\left(\kappa_{\ell} a\right) K_{q}\left(\kappa_{\ell} a\right) \pi
\end{gathered}
$$

and

$$
\widetilde{\mathscr{Y}}_{q}^{(\ell)}=-\epsilon_{q} A_{\ell, \ell} \kappa_{\ell} a K_{q}^{\prime}\left(\kappa_{\ell} a\right) \int_{0}^{\pi}\left(\xi_{\ell}^{P}+\xi_{\ell}^{Q}\right) \cos q \theta \mathrm{~d} \theta .
$$

The integrals $\check{\mathscr{K}}_{q, n}^{(\ell, \ell)}$ and $\widetilde{\mathscr{Y}}_{q}^{(\ell)}$ can be evaluated numerically.
Combining (5.12) and (5.15), we obtain the global matrix equation
$\left[\begin{array}{cccccccc}\mathbb{K}^{(0,0)} & \widehat{\mathbb{K}}^{(0,0)} & \cdots & \mathbb{K}^{(0, \ell)} & {[0]} & \cdots & \mathbb{K}^{\left(0, N_{\xi}\right)} & {[0]} \\ \widetilde{K}^{(0,0)} & \widetilde{\mathbb{K}}^{(0,0)} & \cdots & {[0]} & {[0]} & \cdots & {[0]} & {[0]} \\ \vdots & \vdots & \cdots & \vdots & \vdots & \cdots & \vdots & \vdots \\ \mathbb{K}^{(\ell, 0)} & {[0]} & \cdots & \mathbb{K}^{(\ell, \ell)} & \widehat{\mathbb{K}}^{(\ell, \ell)} & \cdots & \mathbb{K}^{\left(\ell, N_{\xi}\right)} & {[0]} \\ {[0]} & {[0]} & \cdots & \widetilde{K}^{(\ell, \ell)} & \widetilde{\mathbb{K}}^{(\ell, \ell)} & \cdots & {[0]} & {[0]} \\ \vdots & \vdots & \cdots & \vdots & \vdots & \cdots & \vdots & \vdots \\ \mathbb{K}^{\left(N_{\xi}, 0\right)} & {[0]} & \cdots & \mathbb{K}^{\left(N_{\xi}, \ell\right)} & {[0]} & \cdots & \mathbb{K}^{\left(N_{\xi}, N_{\xi}\right)} & \widehat{\mathbb{K}}^{\left(N_{\xi}, N_{\xi}\right)} \\ {[0]} & {[0]} & \cdots & {[0]} & {[0]} & \cdots & \widetilde{\mathbf{K}}^{\left(N_{\xi}, N_{\xi}\right)} & \widetilde{\mathbb{K}}^{\left(N_{\xi}, N_{\xi}\right)}\end{array}\right]\left[\begin{array}{c}\mathscr{X}^{(0)} \\ \widehat{\mathscr{X}}^{(0)} \\ \vdots \\ \mathscr{X}^{(\ell)} \\ \widehat{\mathscr{X}}^{(\ell)} \\ \vdots \\ \mathscr{X}^{\left(N_{\xi}\right)} \\ \widehat{\mathscr{X}}^{\left(N_{\xi}\right)}\end{array}\right]=\left[\begin{array}{c}\mathscr{Y}^{(0)} \\ \widetilde{\mathscr{Y}}^{(0)} \\ \vdots \\ \mathscr{Y}^{(m)} \\ \widetilde{\mathscr{Y}}^{(m)} \\ \vdots \\ \mathscr{Y}^{\left(N_{\xi}\right)} \\ \widetilde{\mathscr{Y}}^{\left(N_{\xi}\right)}\end{array}\right]$
where the unknown vectors are defined by

$$
\mathscr{X}^{(\ell)}=\left[\begin{array}{c}
\xi_{\ell, 1}  \tag{5.17}\\
\vdots \\
\xi_{\ell, n} \\
\vdots \\
\xi_{\ell, N_{P}}
\end{array}\right], \quad \widehat{X}^{(\ell)}=\left[\begin{array}{c}
\widehat{\alpha}_{\ell, 0} \\
\vdots \\
\widehat{\alpha}_{\ell, q} \\
\vdots \\
\widehat{\alpha}_{\ell, N_{\bar{\alpha}}}
\end{array}\right]
$$

and the forcing vectors by

$$
\mathscr{Y}^{(m)}=\left[\begin{array}{c}
\mathscr{Y}_{1}^{(m)}  \tag{5.18}\\
\vdots \\
\mathscr{Y}_{\tau}^{(m)} \\
\vdots \\
\mathscr{Y}_{N_{P}}^{(m)}
\end{array}\right], \quad \widetilde{\mathscr{Y}}^{(m)}=\left[\begin{array}{c}
\widetilde{\mathscr{Y}}_{0}^{(m)} \\
\vdots \\
\widetilde{\mathscr{Y}}_{q}^{(m)} \\
\vdots \\
\widetilde{\mathscr{Y}}_{N_{\tilde{\alpha}}^{(m)}}
\end{array}\right] .
$$

The entries for $\mathbb{K}^{(m, \ell)}, \widehat{\mathbb{K}}^{(m, \ell)}$ and $\check{\mathbb{K}}^{(\ell, \ell)}$ are $\mathscr{K}_{\tau, \ell}^{(m, \ell)}, \widehat{K}_{\tau, p}^{(m, \ell)}$ and $\check{\mathscr{K}}_{q, n}^{(\ell, \ell)}$, respectively. The last matrix $\widetilde{\mathbb{K}}^{(\ell, \ell)}$ is diagonal with entries $\widetilde{K}_{q, q}^{(\ell, \ell)}$. These matrix equations are solved numerically by using the UC Berkeley program Distributed SuperLU which is designed for distributed memory parallel processors, using MPI for interprocess communications. In addition, the Domb-Sykes extrapolation scheme is used to limit the number of evanescent modes $N_{\xi}$.

In the limit of constant depth everywhere, the above matrix equation can be separated into $N_{\xi}+1$ uncoupled matrix equations as follows,

$$
\left[\begin{array}{ll}
\mathbf{K}^{(m, m)} & \widehat{\mathbf{K}}^{(m, m)}  \tag{5.19}\\
\widehat{\mathbf{K}}^{(m, m)} & \widetilde{\mathbf{K}}^{(m, m)}
\end{array}\right]\left[\begin{array}{l}
\mathscr{X}^{(m)} \\
\widehat{\mathscr{X}}^{(m)}
\end{array}\right]=\left[\begin{array}{l}
\mathscr{Y}^{(m)} \\
\widetilde{\mathscr{Y}}^{(m)}
\end{array}\right], \quad m=0,1,2,3, \ldots, N_{\xi} .
$$

Before discussing results from finite-element computations, let us examine a special case of constant depth.

## 6. Semicircular peninsula on constant depth

### 6.1. Analytical solution

As an application, we consider the limiting case of a semi-circular vertical cylinder (peninsula) of radius $a$ attached to a straight coast. The sea depth is constant everywhere. Now there is no need for finite elements; region $\Omega_{F}$ corresponds to the entire sea. The problem can, of course, be solved by combining the method of images and the three-dimensional methods of Kim \& Yue (1989) or of Chau \& Eatock Taylor (1992). In addition to providing information on the effects of incidence angles, this example serves to show that the mild-slope-equation approach is a twodimensional alternative to Chau \& Eatock Taylor. The analytical results are also useful as benchmarks for checking the more numerical tasks involving variable bathymetry.

At first order, the scattered wave coefficient in (4.3) is simply

$$
\begin{equation*}
\alpha_{m}=2 \frac{J_{m}^{\prime}(k a)}{H_{m}^{\prime}(k a)} \cos m \theta_{I}, \quad m=0,1,2,3, \ldots \tag{6.1}
\end{equation*}
$$

At second order, the free wave satisfies the boundary condition on the cylinder

$$
\begin{equation*}
\frac{\partial \xi_{\ell}^{H}}{\partial r}=-\frac{\partial \xi_{\ell}^{P}}{\partial r}, \quad r=a, \quad 0 \leqslant \theta \leqslant \pi, \quad \ell=0,1,2,3 \ldots \tag{6.2}
\end{equation*}
$$

and along the coast.
Using the partial wave expansions for the plane waves, we rewrite $\xi_{\ell}^{P}$ as

$$
\begin{equation*}
\xi_{\ell}^{P}=L_{\ell}^{M} \sum_{m=0}^{\infty} \epsilon_{m} \mathrm{i}^{m} J_{m}(2 k r) 2 \cos m \theta_{I} \cos m \theta+L_{\ell}^{N} \sum_{m=0}^{\infty} \epsilon_{m} \mathrm{i}^{m} J_{m}\left(2 k r \cos \theta_{I}\right) \cos m \theta \tag{6.3}
\end{equation*}
$$

where $L_{\ell}^{M}$ and $L_{\ell}^{N}$ are defined in (4.26). Using (4.23) for $\xi_{\ell}^{H}$ and applying (6.2), we obtain

$$
\begin{equation*}
\widehat{\alpha}_{\ell, m}=-2 \mathrm{i}^{m} L_{\ell}^{M} \frac{J_{m}^{\prime}(2 k a)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} 2 k \cos m \theta_{I}-\mathrm{i}^{m} L_{\ell}^{N} \frac{J_{m}^{\prime}\left(2 k a \cos \theta_{I}\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} 2 k \cos \theta_{I} \tag{6.4}
\end{equation*}
$$

It follows that

$$
\begin{align*}
& \xi_{\ell}^{H}=\sum_{m=0}^{\infty}-\epsilon_{m} \mathrm{i}^{m} K_{m}\left(\kappa_{\ell} r\right) \cos m \theta 2 k\left[L_{\ell}^{M} \frac{2 J_{m}^{\prime}(2 k a)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} \cos m \theta_{I}\right. \\
&  \tag{6.5}\\
& \left.\qquad \quad+L_{\ell}^{N} \frac{J_{m}^{\prime}\left(2 k a \cos \theta_{I}\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} \cos \theta_{I}\right]
\end{align*}
$$

The corresponding potential $\Phi^{H}$ is

$$
\begin{align*}
\Phi^{H}=\sum_{\ell=0}^{\infty} & \frac{\mathrm{i} g}{2 \omega} \frac{\cos \kappa_{\ell}(z+h)}{\cos \kappa_{\ell} h} \mathrm{e}^{-\mathrm{i} 2 \omega t} \sum_{m=0}^{\infty} \epsilon_{m} \mathrm{i}^{m} K_{m}\left(\kappa_{\ell} r\right) \cos m \theta \\
& \times\left[L_{\ell}^{M} \frac{J_{m}^{\prime}(2 k a)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} 4 k \cos m \theta_{I}+L_{\ell}^{N} \frac{J_{m}^{\prime}\left(2 k a \cos \theta_{I}\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right) \kappa_{\ell}} 2 k \cos \theta_{I}\right]+* . \tag{6.6}
\end{align*}
$$

Together with the analytical formula (4.29) for $\xi_{\ell}^{Q}$, the amplitude of the secondorder displacement $\zeta_{2}^{(2)}(r, \theta, t)(\mathrm{cf}$. (3.13)) is

$$
\begin{align*}
\eta_{2,2}^{(2)}(r, \theta)= & \sum_{\ell=0}^{\infty} \sum_{m=0}^{\infty} \epsilon_{m} \cos m \theta\left\{\mathrm{i}^{m} L_{\ell}^{M} 2 \cos m \theta_{I}\left[J_{m}(2 k r)-\frac{2 k}{\kappa_{\ell}} \frac{J_{m}^{\prime}(2 k a)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r\right)\right]\right. \\
& +\mathrm{i}^{m} L_{\ell}^{N}\left[J_{m}\left(2 k r \cos \theta_{I}\right)-\frac{2 k \cos \theta_{I}}{\widehat{\kappa_{\ell}}} \frac{J_{m}^{\prime}\left(2 k a \cos \theta_{I}\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r\right)\right] \\
& \left.+\int_{a}^{\infty} r_{0} \mathrm{~d} r_{0}\left[-\mathrm{i} \frac{2 \omega \mathscr{Q}_{m}\left(r_{0}\right)}{A_{\ell, \ell}}\right] K_{m}\left(\kappa_{\ell} r_{>}\right)\left[\frac{I_{m}^{\prime}\left(\kappa_{\ell} a\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r_{<}\right)-I_{m}\left(\kappa_{\ell} r_{<}\right)\right]\right\} . \tag{6.7}
\end{align*}
$$

Combined with (2.11), the second-order free-surface height is completely determined.
When the incident wave is parallel to the coastline, the problem is identical to a plane wave scattered by a full cylinder in an open sea. (The first-order incident and reflected waves are in the same direction so that their total amplitude is $2 A$.) This closed-form solution should be equivalent to the result of Chau \& Eatock Taylor (1992). As a check for correctness and accuracy, we have compared our numerical results with earlier works for the case of $\theta_{I}=0, r / a=1, h / a=1, \omega^{2} a / g=2$, as plotted in figure 3 by Chau \& Eatock Taylor (1992). The agreement is excellent (Chen 2005).

| $m, n$ | $k a=0.1$ | $k a=1$ | $k a=0.5$ | $k a=10$ |
| :---: | :---: | :---: | :---: | :---: |
| $\theta_{I}=\pi$ | 6,7 | 10,11 | 23,19 | 34,26 |
| $\theta_{I}=5 \pi / 4$ | 6,7 | 10,11 | 22,19 | 34,27 |
| $\theta_{I}=3 \pi / 2$ | 5,6 | 10,10 | 21,18 | 35,27 |

Table 1. Number of terms $(m, n)$ used in the double series for 2.


Figure 2. Dimensionless amplitudes of first-order $|\eta / A|$ along the semi-circular peninsula for three angles of incidence $\theta_{I}=\pi(-), 5 \pi / 4(--), 3 \pi / 2(---)$, and four cylinder radii: (a) $k a=0.1$, (b) $k a=1$, (c) $k a=5$ and (d) $k a=10$. Depth is fixed $k h=1$.

### 6.2. Effects of incidence angle and cylinder radius

We have carried out a number of computations to study the effects of incidence angle and the cylinder radius. For brevity only the free-surface displacements (run-up) along the circumference of the cylinder are shown for a fixed depth-to-wavelength ratio $k h=1$. Three angle of incidences : $\theta_{I}=\pi$ (glancing), $5 \pi / 4$ (oblique) and $3 \pi / 2$ (normal) and four radius-to-wavelength ratios are considered.

In computing the double series of $\mathscr{2}$ defined in (4.14) and (4.15), both $m$ and $n$ must be truncated after a finite number of terms. These numbers are found by numerical experiments for a fixed error allowance, chosen to be $10^{-5}$. As shown in table 1 , the numbers increase for larger $k a$. In addition, the integral for the propagating mode $\ell=0$ in (6.7) is evaluated by sectioning the integration path. Over one section from $a$ to a finite but large $R$, numerical integration is performed by the Gaussian three-points adaptive method. In the remaining section from $R$ to $\infty$, asymptotic approximations are used, as in Chau \& Eatock Taylor (1992).

To help understand the second-order results, we first display in figure 2 the firstorder run-up along the cylinder. As is well known, for the smallest cylinder $k a=0.1$,


Figure 3. Dimensionless amplitudes of second-order $\eta_{2,0}^{(1)} /\left(4 k A^{2}\right)$ along the semicircular peninsula, for different angle of incidence $\theta_{I}=\pi(-), 5 \pi / 4(--), 3 \pi / 2(---)$, and four cylinder radii: (a) $k a=0.1$, (b) $k a=1$, (c) $k a=5$ and (d) $k a=10$. Depth is fixed at $k h=1$.
the run-up is uniform in all directions, and the effect of incidence angle is small. As the cylinder radius increases relative to the wavelength, the effect of incidence angle becomes more pronounced. For glancing incidence, the runup is the greatest on the incidence side (small $\theta$ ), and the smallest in the shadow $(\theta \approx \pi)$. For normal incidence, the run-up varies strongly in different directions with the greatest run-up at the corners $(\theta=0, \pi)$. The directional variation is oscillatory and symmetrical about the shore-normal axis. With increasing $k a$, oscillations become more prominent.

The second-order set-up and set-down along the semicircle is shown in figure 3. This quantity is the time-averaged part of the Bernoulli effect on the free surface (cf. (2.10) and (3.10)). For the smallest cylinder, the maximum set-down occurs in the direction $\theta=\pi / 2$, being the largest for glancing incidence and smallest for normal incidence. With increasing $k a$, the set-up and set-down vary with direction in ways similar to the first-order first-harmonic run-up. The largest mean set-up occurs near $\theta=0$.

The second-harmonic run-up contains two parts (cf. (2.10)). The part due to the firstorder interations (cf. (3.12) is shown in figure 4. For glancing incidence, the greatest run-up is at $\theta=\pi / 2$ for the smallest cylinder, but near the corner facing the incoming wave for the largest cylinder. For normal incidence, the run-up is oscillatory in $\theta$ and symmetrical with respect to the axis $\theta=\pi / 2$, similar to the first-order first-harmonic run-up. The part due to second-order potential is shown in figure 5. For glancing incidence, the greatest response appears at $\theta=0$ for the smallest cylinder, but near


Figure 4. As figure 3, but for $\left|\eta_{2,2}^{(1)}\right| /\left(2 k A^{2}\right)$.


Figure 5. As figure 3, but for $\left|\eta_{2,2}^{(2)}\right| /\left(2 k A^{2}\right)$.


Figure 6. As figure 3, but for $\left|\eta_{2,2}^{(2)}+\eta_{2,2}^{(1)}\right| /\left(2 k A^{2}\right)$.
$\theta=\pi / 2$ for the larger cylinders. The total second-harmonic run-up is also shown in figure 6. The relative magnitudes and phases between $\eta_{2,2}^{(1)}$ and $\eta_{2,2}^{(2)}$ depend on $\theta$. The magnitudes of two parts in the total second-harmonic run-up are comparable.

## 7. A cylinder on top of a semicircular shoal

We now apply the theory for an example combining diffraction and refraction. Consider a semicircular cylinder with radius 20 m standing at the centre of a semicircular shoal next to a straight and cliff-like coast. The top of the shoal is a flat semicircle of radius 20 m at depth 20 m . The sea depth increases monotonically with $r$ until $r=300 \mathrm{~m}$, outside of which the surrounding sea has the greater constant depth of 40 m . The radial variation of the sea depth is given by

$$
h= \begin{cases}20 \mathrm{~m}, & r \leqslant 20 \mathrm{~m}  \tag{7.1}\\ 30-10 \cos \left[\frac{\pi}{280}(r-20)\right] \mathrm{m}, & 20 \mathrm{~m} \leqslant r \leqslant 300 \mathrm{~m} .\end{cases}
$$

The frequency is chosen to be $\omega=0.6873 \mathrm{rads}^{-1}$ such that $k h=2$ in the open sea where $h=40 \mathrm{~m}$. At depth $h=20 \mathrm{~m}$ the local wavenumber is increased to $k h=1.1697$. Two incidence angles ( $\theta_{I}=\pi$ (glancing) and $3 \pi / 2$ (normal)) are considered.

In the hybrid-element scheme, finite elements are used to discretize the shoal, $r<a=300 \mathrm{~m}, 0<\theta<\pi$, outside which ( $\Omega_{F}: r>300 \mathrm{~m}, 0<\theta<\pi$ ) the solution is analytical. Numerical accuracy of the first-order computations depends on the number of finite-element nodes $N_{P}$ (or the grid size), and the number of angular modes $N_{\alpha}$ included in the scattered waves. We have tested two finite-element grids: coarse grid with $N_{P}=73040$ and fine grid with $N_{P}=175625$. Let $k_{0}=0.05 \mathrm{~m}^{-1}$ denote


Figure 7. The first-order amplitude, $|\eta| / A$, over a semicircular shoal around a cylinder. (a) Glancing incidence. (b) Normal incidence.
the incident wavenumber at $h=40 \mathrm{~m}$. The maximum element size $L_{e}$ is such that $k_{0} L_{e}=0.08$ for the coarse grid, and 0.05 for the fine grid. The corresponding ratios of $L_{e}$ to wavelength are quite small: 0.012 and 0.008 , respectively. For either grid, the solution converges to six decimal places if the number of angular modes is $N_{\alpha}=21$ or more. Using 201 angular modes in the outer region, the amplitude for normal incidence at one sample point $\eta(x=0 \mathrm{~m}, y=20 \mathrm{~m}) / A$ is 0.49802 (coarse) and 0.49824 (fine); the relative error is approximately $4 \times 10^{-4}$. Our first-order results are obtained with the fine grid and $N_{\alpha}=201$.

For the second-order problems, we use $N_{P}=175625$ (fine grid) and $N_{\widehat{\alpha}}=91$ (number of angular modes) based on several tests. To limit the number of evanescent modes in the $l$ series, we use the following Domb-Sykes extrapolation scheme. By defining the truncated sum as

$$
\begin{equation*}
y^{(l)}=\sum_{\ell=0}^{l} \xi_{\ell} \tag{7.2}
\end{equation*}
$$

By plotting $y^{(l)}$ vs. $1 / l$, the final result is obtained from the limit of $1 / l=0$. Choosing just two values $l$ and $l+1$, the limit by linear extrapolation gives

$$
\begin{equation*}
y^{(\infty)}=y^{(l+1)}-l\left[y^{(l)}-y^{(l+1)}\right] . \tag{7.3}
\end{equation*}
$$



Figure 8. The second-order set-down/set-up, $\left|\eta_{2,0}^{(1)}\right| / 4 k_{0} A^{2}$, details as figure 7.
We test the number $N_{\xi}$ of evanescent modes required, based on the numerical results with $l=N_{\xi}-1$ and $l+1=N_{\xi}$ in the Domb-Sykes method to extrapolate $y^{(\infty)}$. At the sample point $(x=20 \mathrm{~m}, y=0 \mathrm{~m})$ the numerical results for $\eta_{2,2}^{(2)} / 2 k A^{2}$ are found to be 1.9060 for $N_{\xi}=3$ and 1.9117 for $N_{\xi}=4$. The complex computer algorithm has been checked satisfactorily with the analytical solution of $\S 6$ by using a ring of finite elements surrounding the peninsula (Chen 2005).

We now discuss the numerical results displayed below only for the shoaling region: $r \leqslant a=300 \mathrm{~m}, 0<\theta<\pi$.

The first-order free-surface amplitudes $|\eta| / A$ for glancing and normal incidences are shown in figure 7. For glancing incidence, the problem is equivalent to a full cylinder on a circular shoal in the open sea, attacked by an incident wave of amplitude $2 A$ propagating from left $(x \sim \infty)$ to right $(x \sim-\infty)$. If the cylinder were absent, the incident rays would enter the shoal first and then bend toward the centre of the shoal; the amplitude would grow slowly towards the orgin. On the lee side, rays would be reflected from the coast and intersect with those incident rays farther away from the coast, resulting in constructive interference and greater amplitude (these features are seen from computed results omitted here). With a cylinder of fairly large radius $k_{0} a=1$ where $k_{0}$ is the wavenumber at $h=40 \mathrm{~m}$, strong back-scattering results, as indicated by the wavy envelope in figure 7(a). On the shadow side, only a mild increase


Figure 9. The second-harmonic amplitude computed from the first-harmonic, $\left|\eta_{2,2}^{(1)}\right| / 2 k_{0} A^{2}$, details as figure 7.
with $r$ is seen owing to the constructive inteference between incident and reflected rays which are not intercepted by the cylinder. All variations gradually diminish far outside the shoal not shown here. For normal incidence without the cylinder, the plot of $|\eta / A|$ would resemble that of a simple standing wave with nodal and antinodal lines parallel to the $x$-axis. Now scattering from the cylinder adds modulations in all directions as is evident in figure $7(b)$.

The second-order set-up/set-down of the mean sea level is shown in figure 8 for two incidence angles. Depending solely on the first-order result, the qualitative features resemble those of figure 7. For reference, we recall that the mean sea-level of a progressive wave of amplitude $2 A$ on a sea of constant depth is negative and hence a pure set-down $\eta_{20}^{(1)}=-4 k A^{2} / 2 \sinh 2 k h$. On the other hand, the mean sea-level under a simple standing wave over a horizontal seabed is,

$$
\begin{align*}
\eta_{20} & =k A^{2}\left[\frac{\omega^{2}}{g k} \cos ^{2} k y-\frac{g k}{\omega^{2}} \sin ^{2} k y\right] \\
& =\frac{k A^{2}}{2} \frac{1}{\tanh k h}\left[-\left(1-\tanh ^{2} k h\right)+\cos 2 k y\left(1+\tanh ^{2} k h\right)\right], \tag{7.4}
\end{align*}
$$



Figure 10. The second-order amplitude computed from the second-harmonic, $\left|\eta_{2,2}^{(2)}\right| / 2 k_{0} A^{2}$, details as figure 7 .
which follows from (3.11) by taking $\eta=A \cos k y$. Unlike the simple progressive, wave, the maximum set-up occurs along the lines $2 k y=0,2 \pi, 4 \pi, \ldots$ where $|\eta|=A|\cos k y|=A$ is also the greatest. The maximum set-down occurs at $2 k x=\pi, 3 \pi, 5 \pi, \ldots$, where $|\eta|=A|\cos k y|=0$ is the smallest. These features are qualitatively preserved in the numerical results, as can be seen from figure $8(b)$. In particular, the mean sea-level along the coast is positive everywhere.
The second-order second-harmonic amplitude has two parts. Figure 9 shows for two incidence angles, $\left|\eta_{2,2}^{(1)}\right| / 2 k_{0} A^{2}$ computed from the first-order potential $\Phi_{1}$. The corresponding maximum amplitude of $\zeta_{2}^{(1)}$ is $2\left|\eta_{2,2}^{(1)}\right|$, hence $2 k_{0} A^{2}$ is used as the normalizing scale. The variations again resemble those of the first-order amplitude.

Completing the second-order solution is the part $\eta_{2,2}^{(2)}$ associated with $\Phi_{2}$, shown in figures 10 . Note that, for both incidence angles, the spatial undulations are much more rapid, because the characteristic wavenumber is now $\widehat{\kappa}_{0}$ which is nearly four times the magnitude of $k$ (see (3.21)). As a consequence, the magnitude of the sum of the two complex amplitudes $\left|\eta_{2,2}^{(1)}+\eta_{2,2}^{(2)}\right|$ oscillates nearly twice as fast in space, as is shown in figure 11 after accounting for the phases of the two components.

Finally, we compare the free-surface heights along the circumference of two cylinders of the same radius, $r=r_{a}=20 \mathrm{~m}$ : one on a sea of constant depth ( 40 m )


Figure 11. The total second-order amplitude, $\left|\eta_{2,2}^{(1)}+\eta_{2,2}^{(2)}\right| / 2 k_{0} A^{2}$, details as figure 7.
according to $\S 6$, and the other on top of a shoal just discussed. The incident wave has the same frequency $\omega=0.6873 \mathrm{rad} \mathrm{s}^{-1}$ so that $k_{0} h_{0}=2$ in the sea. Note first by comparing figure $12(a)-12(d)$ with figures $2(b), 3(b), 4(b), 5(b)$ and $6(b)$ that the results on the cylinder for two different constant depths $\left(k_{0} h_{0}=1,2\right)$ are qualtitatively similar; this is probably because the value of $k_{0} r_{a}=1$ is the same in both examples. Secondly, for both incidence angles, shoaling tends to increase mildly the first-order amplitude $|\eta|$ as well as $\left|\eta_{2,2}^{(1)}\right|$, but to decrease the mean sea-level $\left|\eta_{2,0}^{(1)}\right|$, all of which depend solely on the first order. However, the part $\left|\eta_{2,2}^{(2)}\right|$ is considerably more enhanced by shoaling.

## 8. Concluding remarks

We have extended the ideas behind the mild-slope equation, developed previously for linearized problems of water-wave diffraction and refraction, to account for nonlinearity up to second order in wave steepness. In the case of uniform depth, the second-order diffraction is governed by a set of uncoupled two-dimensional Helmholtz equations with forcing. An analytical solution is given for a semi-circular peninsula. A hybrid-element method is described for the second-order refraction/diffraction over a slowly varying bathymetry. Numerical results are demonstrated for the diffraction of a cylinder resting on top of a semi-circular shoal. Applications have been made


Figure 12. Free-surface heights along the semi-circular cylinder of the same radius attacked by an incident wave of the same frequency. For incident wave in deeper water : $k_{0} h_{0}=2$. Thick lines are for variable depth, $h=h(r)$ varying from 40 m to 20 m . Thin lines are for constant depth $h=h_{0}=40 \mathrm{~m}$. Solid lines ( - ): glancing incidence ( $\theta_{I}=\pi$ ). Dashed lines ( --- ): normal incidence $\left(\theta_{I}=3 \pi / 2\right.$ ). (a) $|\eta| / A$, (b) $\left|\eta_{2,0}^{(1)}\right| /\left(4 k_{0} A^{2}\right)$, (c) $\left|\eta_{2,2}^{(1)}\right| /\left(2 k_{0} A^{2}\right)$, (d) $\left|\eta_{2,2}^{(2)}\right| /\left(2 k_{0} A^{2}\right)$.
to other geometries such as pure shoaling without a scatterer and a harbour behind a shoal (Chen 2005). In past studies, the phenomenon of ringing near an offshore tower has been attributed to third-order effects (Faltinsen, Newman \& Vinje 1995). The ideas here can, in principle, be extended to third-order analysis for a tower on a slowly varying seabed, with of course increased algebraic and numerical complexity. Nevertheless, the advantage of this approach that discrete computations are required only for two horizontal coordinates, would probably be even more preferable since refraction usually involves a large domain extending over many wavelengths in all horizontal directions. The fully three-dimensional alternative via either boundary elements or finite elements covering the entire region of variable boundary and bathymetry would appear to be extremely cumbersome and demanding. On the other hand, extending the three-dimensional treatment of Yue et al. (1978) and Kim \& Yue (1989) only near the structure, and matching with the quasi-two-dimensional mild-slope approximation for the much larger zone of variable depth, may simplify the numerical task considerably.

We point out that the present theory is an extension of Stokes approximation, which is inherently limited to a small Ursell-Stokes parameter $\left(A / k^{2} h^{3} \ll 1\right)$. For very
shallow water waves propagating over long distances, we must use Boussinesq or other appropriate approximations expressly devised for small $k h$.

Finally, for treating slow-drift motions of floating platforms or long-period oscillations in harbours, it is necessary to consider nonlinear effects due to random incident waves of broad frequency band. The present theory can be extended to facilitate the analysis, as will be reported in the near future.

The authors thank the US Office of Naval Research (Grant N00014-04-1-0077), US National Science Foundation (Grant CTS-0075713) and the US-Israel Binational Science Foundation (Grant 200405) for their financial support of this study.

## Appendix A. Derivation of second-order mild-slope equation

We outline the derivation by following the procedure of Smith \& Sprinks (1975) for the linearized homogeneous problem with the frequency $\omega$. Except for the inhomogeneous term, and the change of frequency to $2 \omega$, the result is similar to that of Porter \& Staziker (1995) derived by a different method for the first-order problem.

The function $\psi(x, y, z)$ satisfies

$$
\begin{gather*}
\nabla^{2} \psi+\frac{\partial^{2} \psi}{\partial z^{2}}=0, \quad-h(x, y)<z<0  \tag{A1}\\
\frac{\partial \psi}{\partial z}=-\nabla \psi \cdot \nabla h, \quad z=-h(x, y)  \tag{A2}\\
\frac{\partial \psi}{\partial z}-\frac{4 \omega^{2}}{g} \psi=F, \quad z=0 \tag{A3}
\end{gather*}
$$

Let us introduce

$$
\begin{equation*}
f_{m}=\frac{\cos \kappa_{m}(z+h)}{\cos \kappa_{m} h} \tag{A4}
\end{equation*}
$$

which satisfies

$$
\begin{gather*}
\frac{\partial^{2} f_{m}}{\partial z^{2}}+\kappa_{m}^{2} f_{m}=0, \quad-h<z<0  \tag{A5}\\
\frac{\partial f_{m}}{\partial z}=0, \quad z=-h  \tag{A6}\\
\frac{\partial f_{m}}{\partial z}-\frac{4 \omega^{2}}{g} f_{m}=0, \quad z=0 \tag{A7}
\end{gather*}
$$

where $\kappa_{m}, m=0,1,2, \ldots$ are defined by (3.20) and $\kappa_{o}=-\mathrm{i} \widehat{\kappa}_{0}$ is defined by (3.21).
Regarding (A1) as an ordinary differential equation in $z$ and applying Green's formula to $f_{m}$ and $\psi$, we obtain, after using (A 1)-(A 7),

$$
\begin{equation*}
\int_{-h}^{0}\left(-\psi \kappa_{m}^{2} f_{m}+f_{m} \nabla^{2} \psi\right) \mathrm{d} z=-\left(f_{m}\right)_{z=0} F-\left(f_{m} \nabla \psi \cdot \nabla h\right)_{z=-h} \tag{A8}
\end{equation*}
$$

With the assumption,

$$
\begin{equation*}
\psi=-\frac{\mathrm{i} g}{2 \omega} \sum_{\ell=0}^{\infty} \xi_{\ell}(x, y) f_{\ell}(x, y, z) \tag{A9}
\end{equation*}
$$

(A 8) becomes

$$
\begin{align*}
\sum_{\ell=0}^{\infty} & \left\{-\xi_{\ell} \kappa_{m}^{2} g \int_{-h}^{0} f_{m} f_{\ell} \mathrm{d} z+\xi_{\ell} g \int_{-h}^{0} f_{m} \frac{\partial f_{\ell}}{\partial h} \nabla^{2} \mathrm{~d} z+\xi_{\ell} g \int_{-h}^{0} f_{m} \frac{\partial^{2} f_{\ell}}{\partial h^{2}}(\nabla h)^{2} \mathrm{~d} z\right. \\
& \left.+\left(\nabla^{2} \xi_{\ell}\right) g \int_{-h}^{0} f_{m} f_{\ell} \mathrm{d} z+\left(\nabla \xi_{\ell}\right) \cdot\left(g \int_{-h}^{0} 2 f_{m} \frac{\partial f_{\ell}}{\partial h} \nabla h d z\right)\right\} \\
= & -2 \mathrm{i} \omega F-\sum_{\ell=0}^{\infty}\left\{\left[\xi_{\ell} g f_{m} \frac{\partial f_{\ell}}{\partial h}(\nabla h)^{2}+\left(\nabla \xi_{\ell}\right) \cdot\left(g f_{m} f_{\ell} \nabla h\right)\right]_{z=-h}\right\} \tag{A10}
\end{align*}
$$

After using Leibniz' rule,

$$
\begin{align*}
\nabla \cdot\left(\int_{-h}^{0} f_{m} f_{\ell} \nabla \xi_{\ell} d z\right)= & \int_{-h}^{0} f_{m} f_{\ell} \nabla^{2} \xi_{\ell} \mathrm{d} z+\int_{-h}^{0} f_{m} \frac{\partial f_{\ell}}{\partial h} \nabla h \cdot \nabla \xi_{\ell} \mathrm{d} z \\
& +\int_{-h}^{0} f_{\ell} \frac{\partial f_{m}}{\partial h} \nabla h \cdot \nabla \xi_{\ell} d z+\left(f_{\ell} f_{m} \nabla \xi_{\ell} \cdot \nabla h\right)_{z=-h} \tag{A11}
\end{align*}
$$

(A 10) can be written as

$$
\begin{align*}
\sum_{\ell=0}^{\infty} & \left\{\xi_{\ell}\left(-\kappa_{m}^{2} g \int_{-h}^{0} f_{m} f_{\ell} \mathrm{d} z\right)+\nabla \cdot\left[\left(\nabla \xi_{\ell}\right)\left(g \int_{-h}^{0} f_{m} f_{\ell} \mathrm{d} z\right)\right]\right. \\
& +\int_{-h}^{0} g\left(f_{m} \frac{\partial f_{\ell}}{\partial h}-f_{\ell} \frac{\partial f_{m}}{\partial h}\right) \nabla h \cdot \nabla \xi_{\ell} \mathrm{d} z+\int_{-h}^{0} g f_{m} \frac{\partial f_{\ell}}{\partial h} \xi_{\ell} \nabla^{2} h \\
& \left.+\int_{-h}^{0} g f_{m} \xi_{\ell} \frac{\partial^{2} f_{\ell}}{\partial h^{2}}(\nabla h)^{2} \mathrm{~d} z+g\left[f_{m} \xi_{\ell} \frac{\partial f_{\ell}}{\partial h}(\nabla h)^{2}\right]_{z=-h}\right\}=-\mathrm{i} 2 \omega F \tag{A12}
\end{align*}
$$

Equation (A 12) can be simplified to the form

$$
\begin{equation*}
\sum_{\ell=0}^{\infty}\left\{\nabla \cdot\left(A_{m, \ell} \nabla \xi_{\ell}\right)+B_{m, \ell} \nabla h \cdot \nabla \xi_{\ell}+C_{m, \ell} \xi_{\ell}\right\}=-\mathrm{i} 2 \omega F \tag{A13}
\end{equation*}
$$

where

$$
\begin{gather*}
A_{m, \ell}=g \int_{-h}^{0}\left(f_{m} f_{\ell}\right) \mathrm{d} z  \tag{A14}\\
B_{m, \ell}=g\left(U_{m, \ell}-U_{\ell, m}\right)  \tag{A15}\\
U_{m, \ell}=\int_{-h}^{0}\left(f_{m} \frac{\partial f_{\ell}}{\partial h}\right) \mathrm{d} z  \tag{A16}\\
C_{m, \ell}=-\kappa_{m}^{2} A_{m, \ell}+g U_{m, \ell} \nabla^{2} h+g V_{m, \ell}(\nabla h)^{2} \tag{A17}
\end{gather*}
$$

and

$$
\begin{equation*}
V_{m, \ell}=\int_{-h}^{0}\left(f_{m} \frac{\partial^{2} f_{\ell}}{\partial h^{2}}\right) \mathrm{d} z+\left[f_{m} \frac{\partial f_{\ell}}{\partial h}\right]_{z=-h} \tag{A18}
\end{equation*}
$$

Using Leibniz' rule, we can rewrite $V_{m, \ell}$ as

$$
\begin{align*}
V_{m, \ell} & =\frac{\partial}{\partial h} \int_{-h}^{0}\left(f_{m} \frac{\partial f_{\ell}}{\partial h}\right) \mathrm{d} z-\int_{-h}^{0}\left(\frac{\partial f_{m}}{\partial h} \frac{\partial f_{\ell}}{\partial h}\right) \mathrm{d} z \\
& =\frac{\partial U_{m, \ell}}{\partial h}-\int_{-h}^{0}\left(\frac{\partial f_{m}}{\partial h} \frac{\partial f_{\ell}}{\partial h}\right) \mathrm{d} z \tag{A19}
\end{align*}
$$

By using the orthogonality of the vertical eigenfunctions $\left\{f_{m}\right\}, A_{m, \ell}$ is found as given by (3.23). After straightforward but lengthy algebra (Chen 2005), we obtain the coefficients given explicitly by (3.25) to (3.28).

## Appendix B. Equivalence to Stokes waves

We show here that (4.27) and (4.28) are equivalent to Stokes waves of the following standard form:

$$
\begin{align*}
\binom{\Phi_{2}^{I}}{\Phi_{2}^{R}}_{\text {Stokes }}= & -\frac{3 \mathrm{i} \omega}{4} \frac{\cosh 2 k(z+h)}{\sinh ^{4} k h} \frac{A^{2}}{4} \exp \left(\mathrm{i} 2 k r \cos \left(\theta \mp \theta_{I}\right)-2 \mathrm{i} \omega t\right)+*  \tag{B1}\\
\left(\Phi_{2}^{I R}\right)_{\text {Stokes }}= & -\frac{\mathrm{i} A^{2} \omega}{2 g} \frac{1+2 \cos 2 \theta_{I}-3 \tanh ^{2} k h}{\tanh k h\left[2 \cos \theta_{I} \tanh \left(2 k h \cos \theta_{I}\right)-4 \tanh k h\right]} \\
& \times \frac{\cosh \left[2 k(z+h) \cos \theta_{I}\right]}{\cosh \left(2 k h \cos \theta_{I}\right)} \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta-2 \mathrm{i} \omega t\right)+* \tag{B2}
\end{align*}
$$

To show this, we first obtain from (3.16),

$$
\begin{equation*}
\widehat{\beta}-\bar{\beta} k^{2}=\frac{3 \omega k}{\sinh k h \cosh k h}, \quad \widehat{\beta}-\bar{\beta} k^{2} \cos \left(2 \theta_{I}\right)=\frac{\mathrm{i} k \omega}{\tanh k h}\left[1+2 \cos \left(2 \theta_{I}\right)-3 \tanh ^{2} k h\right] \tag{B3}
\end{equation*}
$$

The Stokes wave potential can be rewritten as

$$
\begin{align*}
&\binom{\Phi_{2}^{I}}{\Phi_{2}^{R}}_{\text {Stokes }}=-\left[\frac{\cosh k h}{4 k \sinh ^{3} k h} \cosh 2 k(z+h)\right] \\
& \times\left(\widehat{\beta}-\bar{\beta} k^{2}\right) \frac{A^{2}}{4} \exp \left(\mathrm{i} 2 k r \cos \left(\theta-\theta_{I}\right)-2 \mathrm{i} \omega t\right)+*  \tag{B4}\\
&\left(\Phi_{2}^{I R}\right)_{\text {Stokes }}=-\frac{\mathrm{i} A^{2} \omega}{2 g} \frac{\widehat{\beta}-\bar{\beta} k^{2} \cos \left(2 \theta_{I}\right)}{\left[2 \cos \theta_{I} \tanh \left(2 k h \cos \theta_{I}\right)-4 \tanh k h\right]} \\
& \times \frac{\cosh \left[2 k(z+h) \cos \theta_{I}\right]}{\cosh \left(2 k h \cos \theta_{I}\right)} \exp \left(\mathrm{i} 2 k r \cos \theta_{I} \cos \theta-2 i \omega t\right)+* \tag{B5}
\end{align*}
$$

Let us expand $\cosh 2 k(z+h)$ and $\cosh \left[2 k(z+h) \cos \theta_{I}\right]$ in terms of the vertical eigenfunctions

$$
\begin{equation*}
\cosh 2 k(z+h)=\sum_{\ell=0}^{\infty} a_{\ell} f_{\ell}, \quad \cosh \left[2 k(z+h) \cos \theta_{I}\right]=\sum_{\ell=0}^{\infty} b_{\ell} f_{\ell} \tag{B6}
\end{equation*}
$$

where

$$
\begin{equation*}
f_{\ell}=\frac{\cos k_{\ell}(z+h)}{\cos \kappa_{\ell} h} \tag{B7}
\end{equation*}
$$

By using the orthogonality of $f_{\ell}$, the dispersion relation as well as the identities

$$
\begin{equation*}
2 \tanh 2 k h-4 \tanh k h=-\frac{4 \sinh ^{3} k h}{\cosh k h \cosh 2 k h}, \quad \kappa_{\ell} \tan \kappa_{\ell} h=-\frac{4 \omega^{2}}{g}=-4 k \tanh k h \tag{B8}
\end{equation*}
$$

we obtain

$$
\begin{equation*}
a_{\ell}=-\frac{g}{A_{\ell, \ell}} \frac{4 k \sinh ^{3} k h}{\cosh k h} \frac{1}{4 k^{2}+\kappa_{\ell}^{2}} \tag{B9}
\end{equation*}
$$

$$
\begin{equation*}
b_{\ell}=-\frac{g}{A_{\ell, \ell}} \frac{\left[2 k \cos \theta_{I} \tanh \left(2 k h \cos \theta_{I}\right)-4 k \tanh k h\right] \cosh \left(2 k h \cos \theta_{I}\right)}{\left(4 k^{2}+\kappa_{\ell}^{2}\right)} \tag{B10}
\end{equation*}
$$

where $A_{\ell, \ell}$ is given by (3.23). The equivalence of (B4) and (4.27) and of (B 5) and (4.28) is evident.

## Appendix C. Two-dimensional Green's function

We define the Green function by

$$
\begin{align*}
\nabla^{2} G_{\ell}-\kappa_{\ell}^{2} G_{\ell}=\frac{1}{r} \delta\left(r-r_{0}\right) & {\left[\delta\left(\theta-\theta_{0}\right)+\delta\left(\theta+\theta_{0}\right)\right], \quad r, r_{0}>a }  \tag{C1}\\
\frac{\partial G_{\ell}}{\partial \theta} & =0, \quad \theta=0, \pi  \tag{C2}\\
\frac{\partial G_{\ell}}{\partial r} & =0, \quad r=a \tag{C3}
\end{align*}
$$

For the propagating mode $\ell=0$, the usual (strong) radiation condition is required,

$$
\begin{equation*}
\sqrt{r}\left(\frac{\partial G_{0}}{\partial r}-\mathrm{i} \widehat{\kappa}_{0} G_{0}\right) \rightarrow 0, \quad r \rightarrow \infty \tag{C4}
\end{equation*}
$$

For the evanescent modes $\ell=1,2,3, \ldots$, we require that

$$
\begin{equation*}
G_{\ell} \rightarrow 0, \quad r \rightarrow \infty, \quad \ell=1,2,3, \ldots \tag{C5}
\end{equation*}
$$

By the standard method of eigenfunction expansions, we obtain the following result,

$$
\begin{equation*}
G_{\ell}=\frac{1}{\pi} \sum_{m=0}^{\infty} \epsilon_{m} g_{\ell m} \cos m \theta \cos m \theta_{0} \tag{C6}
\end{equation*}
$$

where

$$
\begin{equation*}
g_{\ell m}\left(r, r_{0}\right)=\left[\frac{I_{m}^{\prime}\left(\kappa_{\ell} a\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r_{0}\right)-I_{m}\left(\kappa_{\ell} r_{0}\right)\right] K_{m}\left(\kappa_{\ell} r\right), \quad r>r_{0}, \tag{C7}
\end{equation*}
$$

and

$$
\begin{equation*}
g_{\ell m}\left(r, r_{0}\right)=\left[\frac{I_{m}^{\prime}\left(\kappa_{\ell} a\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r\right)-I_{m}\left(\kappa_{\ell} r\right)\right] K_{m}\left(\kappa_{\ell} r_{0}\right), \quad r<r_{0} \tag{C8}
\end{equation*}
$$

Clearly, the Green function is symmetric: $G_{\ell}\left(r, \theta ; r_{0}, \theta_{0}\right)=G_{\ell}\left(r_{0}, \theta_{0} ; r, \theta\right)$. Defining

$$
\begin{equation*}
r_{>}=\operatorname{Max}\left(r, r_{0}\right), \quad r_{<}=\operatorname{Min}\left(r, r_{0}\right) \tag{C9}
\end{equation*}
$$

we express the Green function in a more compact form,

$$
\begin{equation*}
G_{\ell}\left(r_{>}, r_{<}\right)=\sum_{m=0}^{\infty} \frac{\epsilon_{\ell}}{\pi}\left[\frac{I_{m}^{\prime}\left(\kappa_{\ell} a\right)}{K_{m}^{\prime}\left(\kappa_{\ell} a\right)} K_{m}\left(\kappa_{\ell} r_{<}\right)-I_{m}\left(\kappa_{\ell} r_{<}\right)\right] K_{m}\left(\kappa_{\ell} r_{>}\right) \cos m \theta \cos m \theta_{0} \tag{C10}
\end{equation*}
$$

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